



# Introduction to X-ray Absorption Spectroscopy

Slides adapted from materials by Bruce Ravel (NIST & NSLS-II) and Eli Stavitski (NSLS-II)

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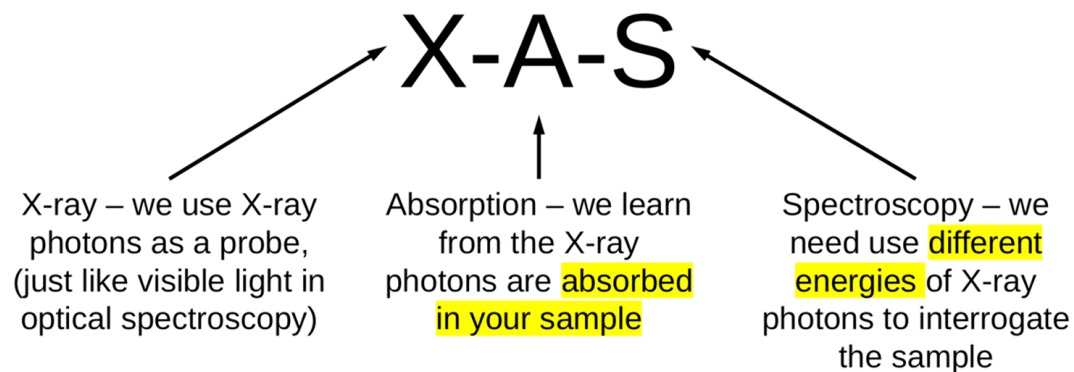
3/10/2026

# Outline

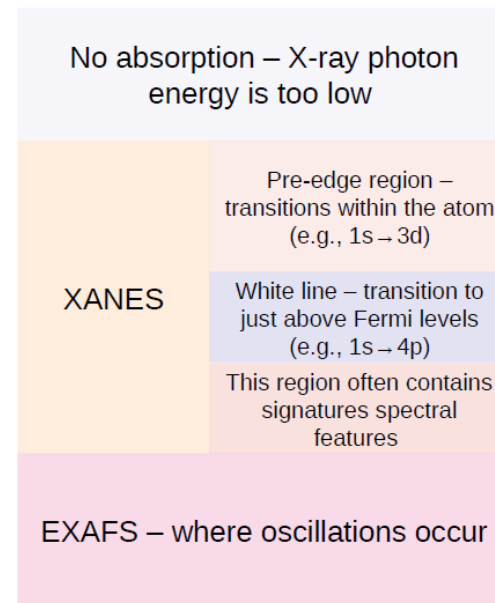
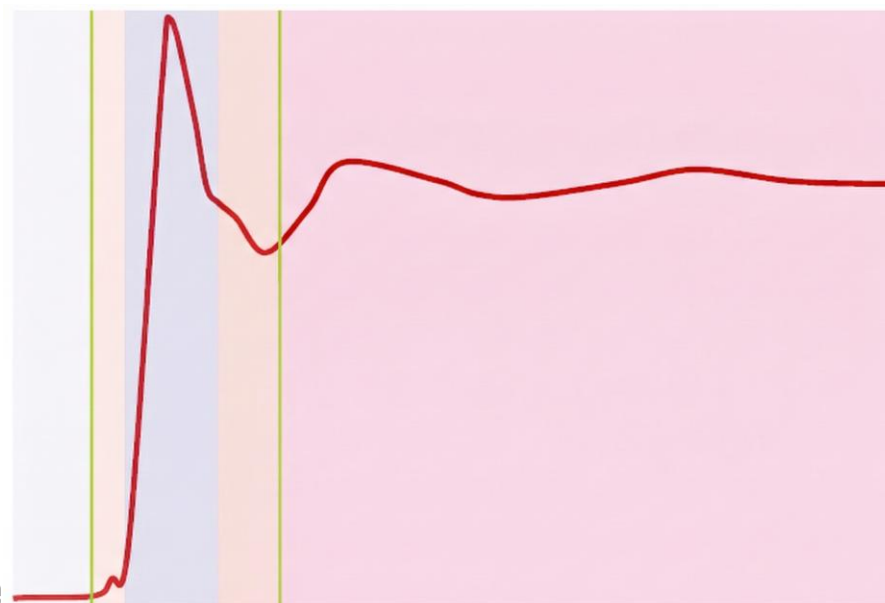
- What is XAS?
- Basic physics of X-ray absorption
- Beamline setup and detection
- What XAS can tell us

# What is XAS?

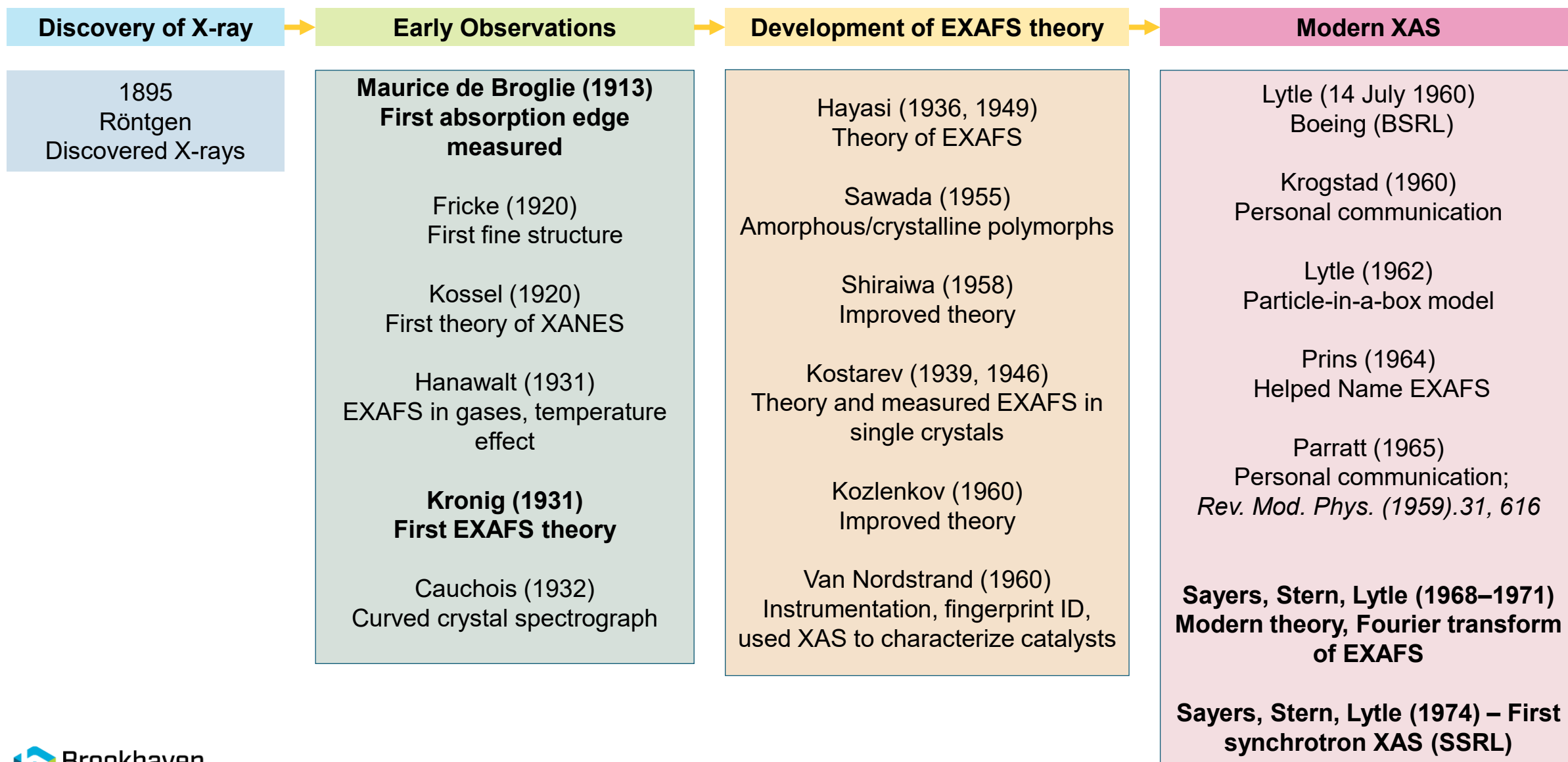
X-ray Absorption Spectroscopy (**XAS**) measures how the X-ray absorption coefficient changes near an absorption edge.



- XAFS** X-ray Absorption Fine-Structure Spectroscopy (= XAS)
- XANES** X-ray Absorption Near-Edge Spectroscopy
- EXAFS** Extended X-ray Absorption Fine-Structure
- NEXAFS** Near Edge X-ray Absorption Fine Structure

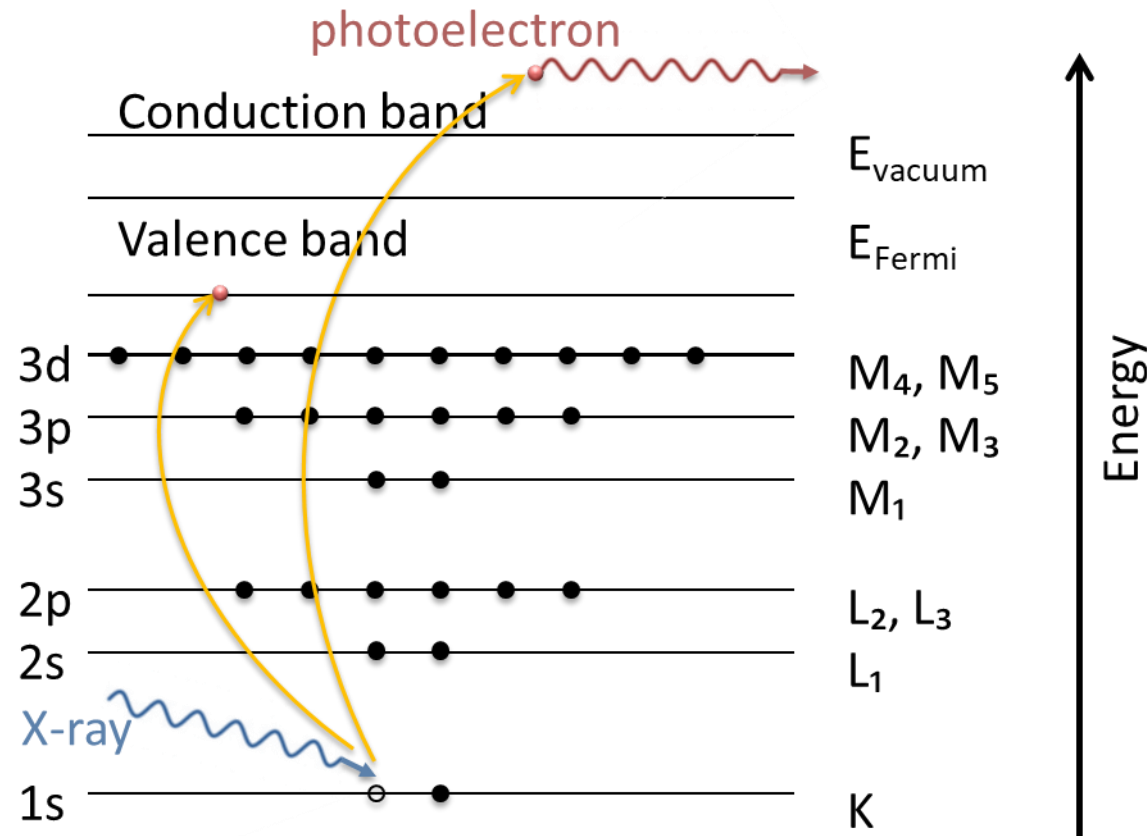


# History of XAS



# Basic physics of XAS

- XAS is an inner-shell spectroscopy because X-rays interact with core electrons (1s, 2s, 2p).
- An X-ray photon excites a core electron into an unoccupied state or into the continuum.
- X-ray absorption occurs through the photoelectric effect.

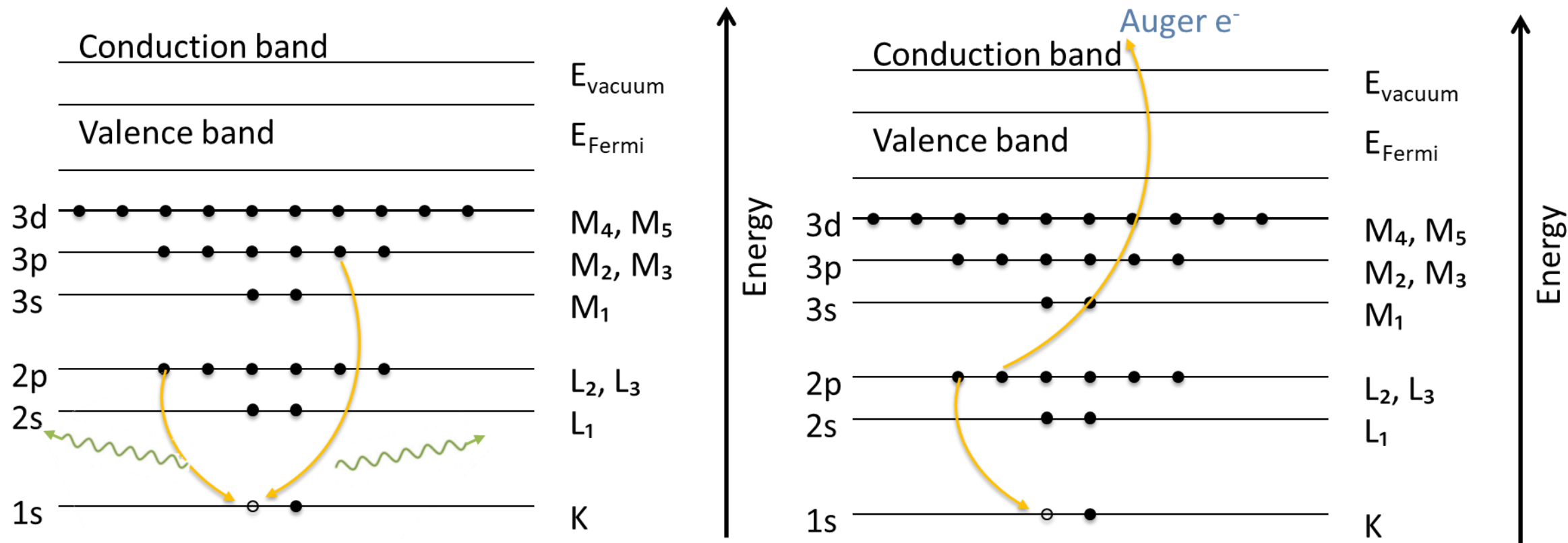


# Basic physics of XAS

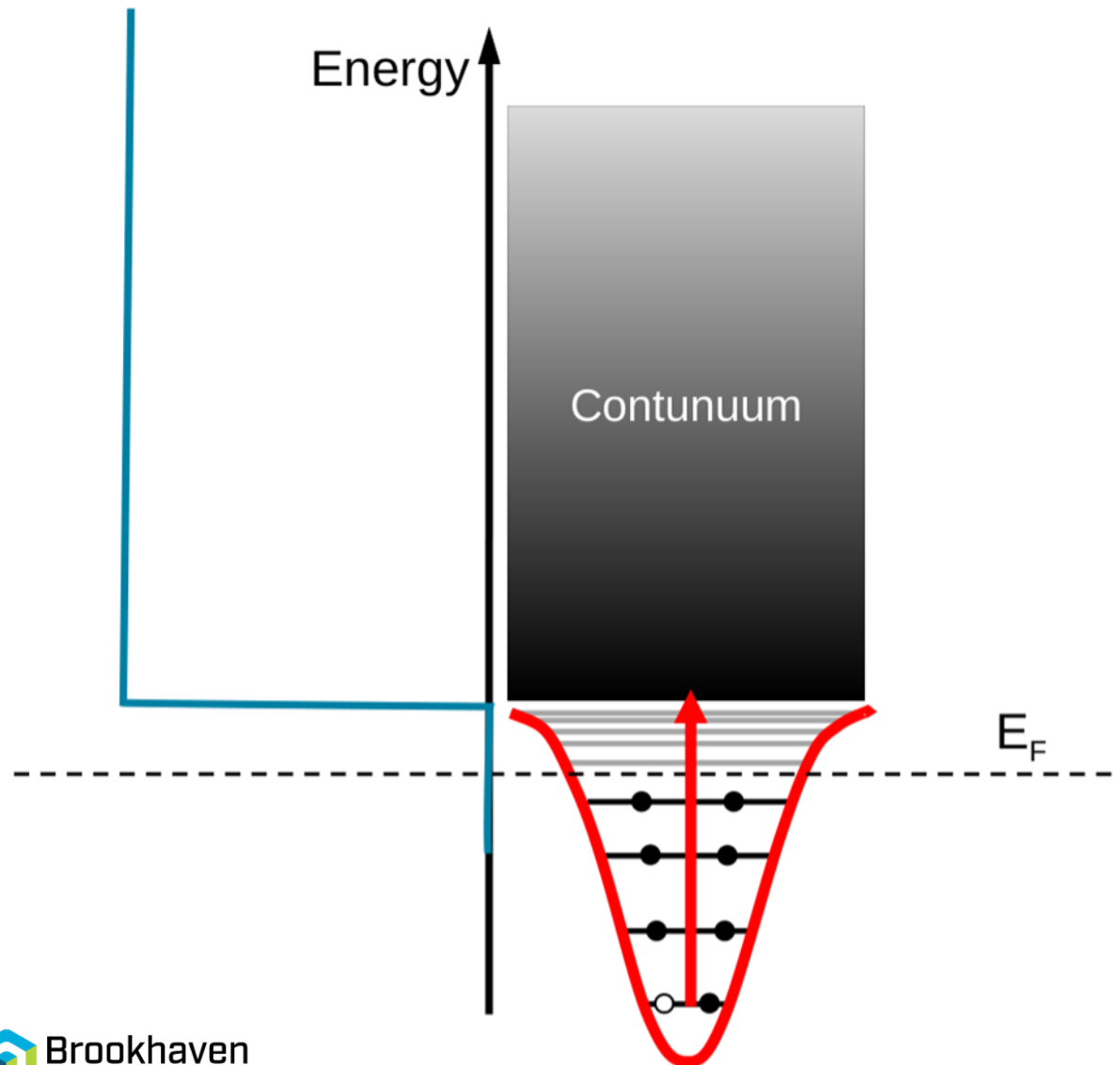
After X-ray absorption, the atom relaxes as a higher-energy electron fills the core hole.

Two common decay pathways:

- Fluorescent X-ray
- Auger electron



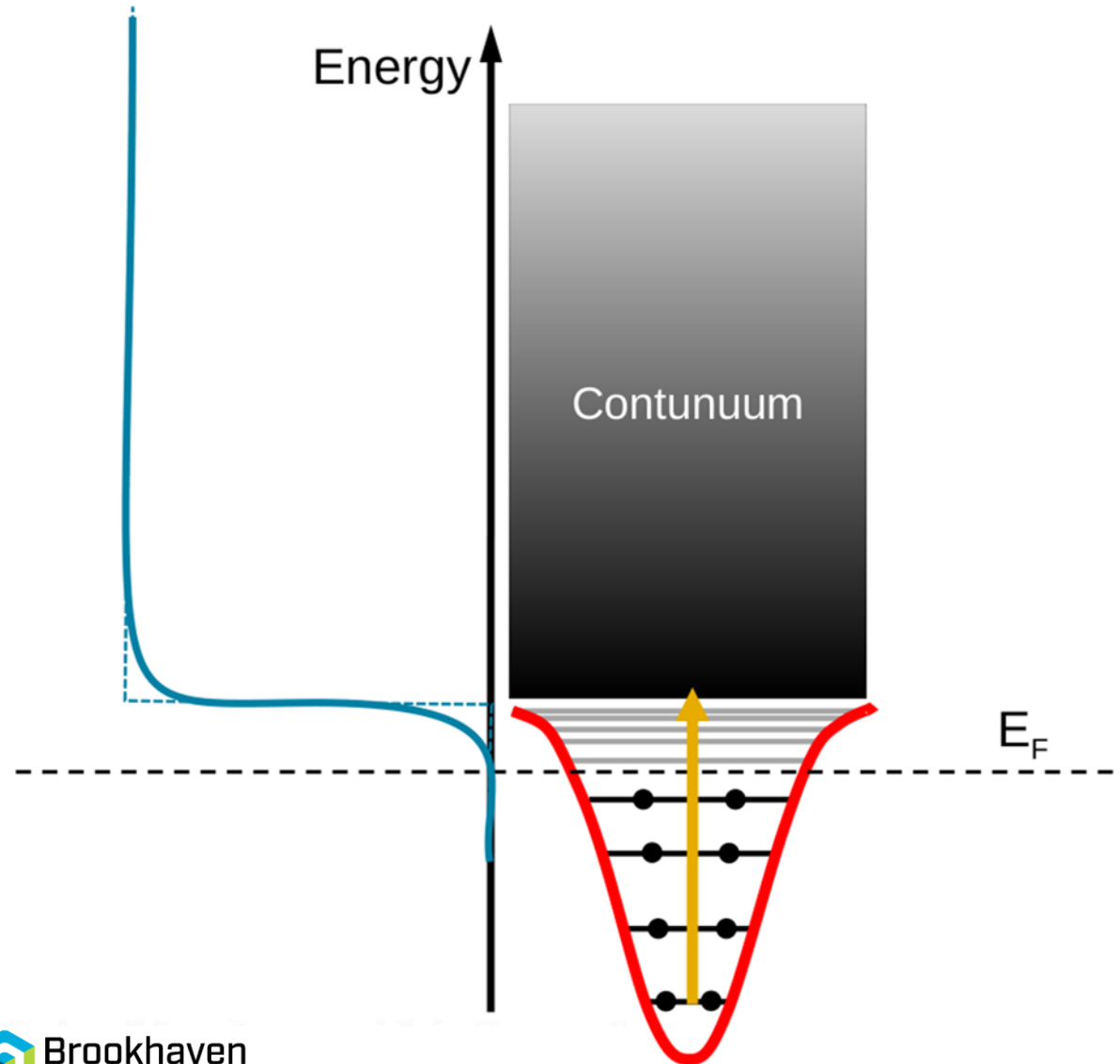
# Processes behind X-ray absorption (1)



Example: K-edge of a 3d element

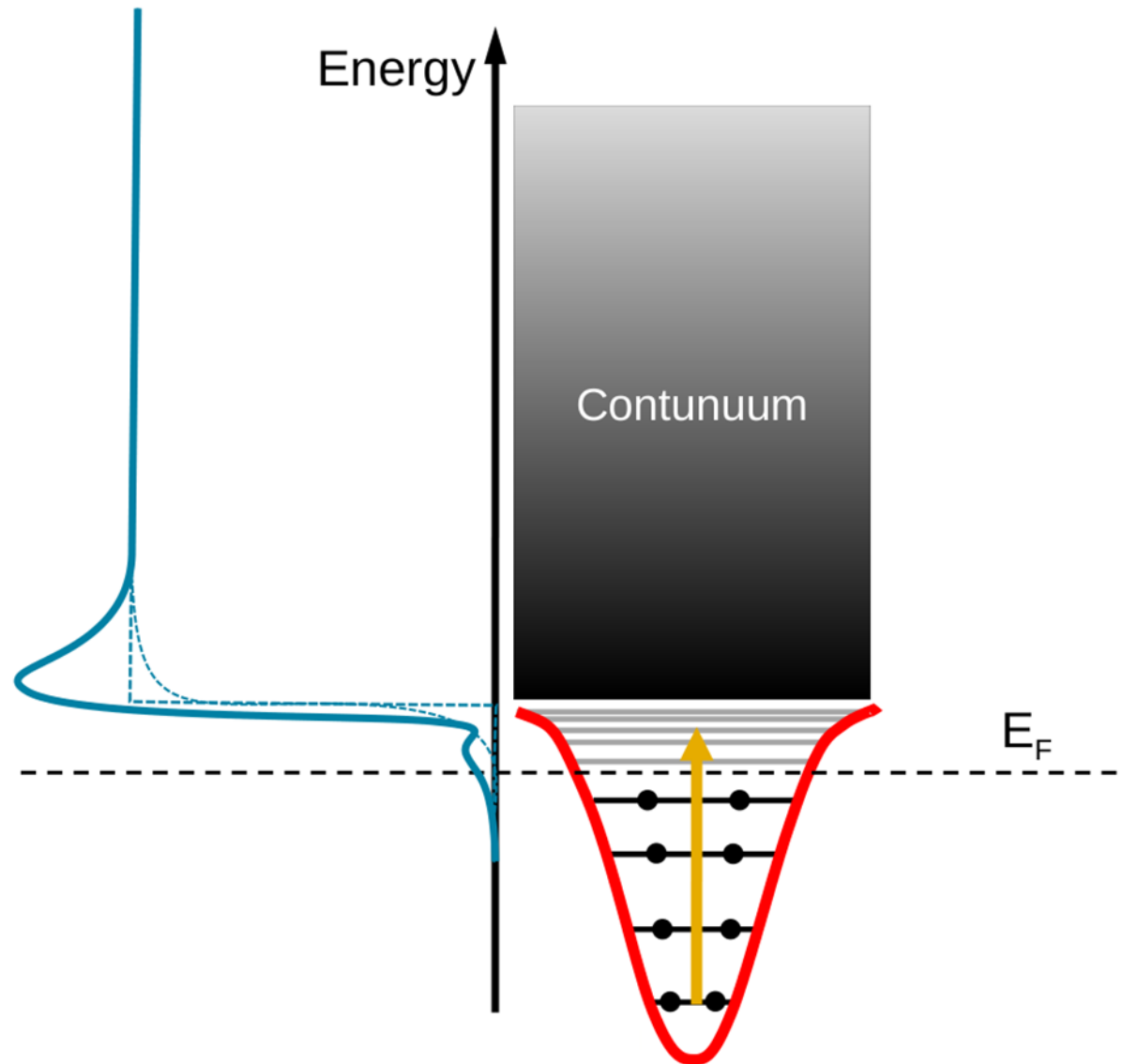
- When the photon energy exceeds the binding energy of a core electron, a photoelectron is emitted.
- This produces the sharp rise in absorption known as the absorption edge.

# Processes behind X-ray absorption (2)



- Once the 1s electron is ejected, the core hole is filled within femtoseconds.
- Because the core-hole lifetime is short, the excited state has an intrinsic energy uncertainty.
- This lifetime broadening is typically on the order of 1 to several eV.

# Processes behind X-ray absorption (3)

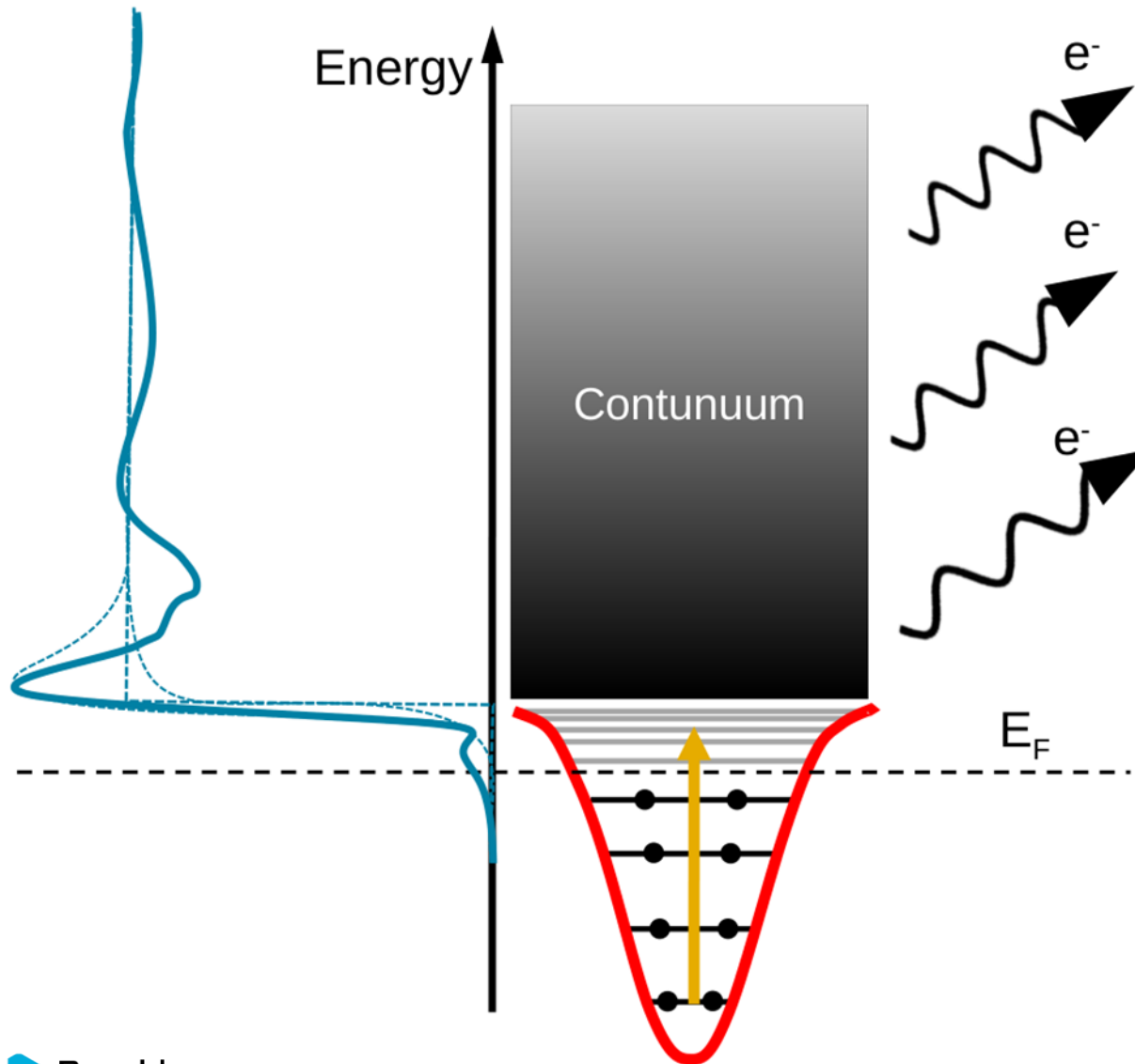


- Now consider the unoccupied states just above the Fermi level.

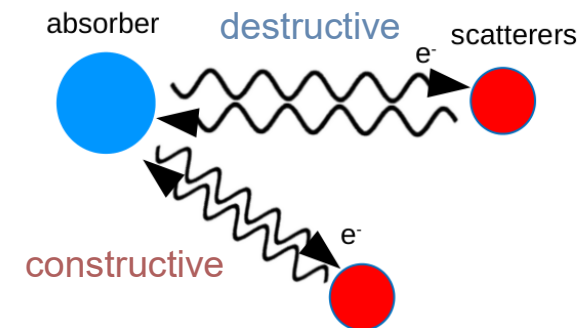


- According to the dipole selection rule,  $1s \rightarrow 4p$  transitions are dipole allowed and therefore give strong absorption.
- $1s \rightarrow 3d$  transitions are dipole forbidden, but weak quadrupole transitions may become observable depending on symmetry.

# Processes behind X-ray absorption (4)



- In condensed matter, the emitted photoelectron is scattered by neighboring atoms.
- The outgoing and back-scattered photoelectron waves interfere.
- This interference modulates the absorption coefficient and gives rise to EXAFS oscillations.



# Selection rules

The absorption probability is determined by the transition matrix element between the initial and final states.

This transition probability can be derived from Fermi's Golden Rule.

$$\mu(E) \propto \sum_f |\langle \psi_f | \hat{\epsilon} \cdot \mathbf{r} e^{i\mathbf{k} \cdot \mathbf{r}} | \psi_i \rangle|^2$$

Dipole selection rule:  $\Delta l = \pm 1$

Initial state:		final state:
K-edge: 1s (n=1, l=0, m=0)	→	p
L1-edge: 2s (n=2, l=0, m=0)	→	p
L2-edge: 2p (j=1/2) (n=2, l=1)	→	d(&s)
L3-edge: 2p (j=3/2) (n=2, l=1)	→	d(&s)

Selection rules for discrete transitions can be found here,  
<https://physics.nist.gov/Pubs/AtSpec/node17.html>

# What elements can we probe with hard X-rays?

**Periodic Table of the Elements**

The periodic table shows the following elements and their atomic weights:

1 IA 1A 1 <b>H</b> Hydrogen 1.008	2 IIA 2A 4 <b>Be</b> Beryllium 9.012											13 IIIA 3A 5 <b>B</b> Boron 10.811	14 IVA 4A 6 <b>C</b> Carbon 12.011	15 VA 5A 7 <b>N</b> Nitrogen 14.007	16 VIA 6A 8 <b>O</b> Oxygen 15.999	17 VIIA 7A 9 <b>F</b> Fluorine 18.998	18 VIIIA 8A 10 <b>Ne</b> Neon 20.180			
3 <b>Li</b> Lithium 6.941	11 <b>Na</b> Sodium 22.990	12 <b>Mg</b> Magnesium 24.305	19 <b>K</b> Potassium 39.098	20 <b>Ca</b> Calcium 40.078	21 <b>Sc</b> Scandium 44.956	22 <b>Ti</b> Titanium 47.88	23 <b>V</b> Vanadium 50.942	24 <b>Cr</b> Chromium 51.996	25 <b>Mn</b> Manganese 54.938	26 <b>Fe</b> Iron 55.933	27 <b>Co</b> Cobalt 58.933	28 <b>Ni</b> Nickel 58.693	29 <b>Cu</b> Copper 63.546	30 <b>Zn</b> Zinc 65.39	31 <b>Ga</b> Gallium 69.732	32 <b>Ge</b> Germanium 72.61	33 <b>As</b> Arsenic 74.922	34 <b>Se</b> Selenium 78.09	35 <b>Br</b> Bromine 79.904	36 <b>Kr</b> Krypton 84.80
37 <b>Rb</b> Rubidium 84.468	38 <b>Sr</b> Strontium 87.62	39 <b>Y</b> Yttrium 88.906	55 <b>Cs</b> Cesium 132.905	56 <b>Ba</b> Barium 137.327	57-71 Lanthanide Series	72 <b>Hf</b> Hafnium 178.49	73 <b>Ta</b> Tantalum 180.948	74 <b>W</b> Tungsten 183.85	75 <b>Re</b> Rhenium 186.207	76 <b>Os</b> Osmium 190.23	77 <b>Ir</b> Iridium 192.22	78 <b>Pt</b> Platinum 195.08	79 <b>Au</b> Gold 196.967	80 <b>Hg</b> Mercury 200.59	81 <b>Tl</b> Thallium 204.383	82 <b>Pb</b> Lead 207.2	83 <b>Bi</b> Bismuth 208.980	84 <b>Po</b> Polonium [208.982]	85 <b>At</b> Astatine 209.987	86 <b>Rn</b> Radon 222.018
87 <b>Fr</b> Francium 223.020	88 <b>Ra</b> Radium 226.025	89-103 Actinide Series	104 <b>Rf</b> Rutherfordium [261]	105 <b>Db</b> Dubnium [262]	106 <b>Sg</b> Seaborgium [266]	107 <b>Bh</b> Bohrium [264]	108 <b>Hs</b> Hassium [269]	109 <b>Mt</b> Meitnerium [268]	110 <b>Ds</b> Darmstadtium [269]	111 <b>Rg</b> Roentgenium [272]	112 <b>Cn</b> Copernicium [277]	113 <b>Uut</b> Ununtrium unknown	114 <b>Fl</b> Flerovium [289]	115 <b>Uup</b> Ununpentium unknown	116 <b>Lv</b> Livermorium [298]	117 <b>Uus</b> Ununseptium unknown	118 <b>Uuo</b> Ununoctium unknown			
			57 <b>La</b> Lanthanum 138.906	58 <b>Ce</b> Cerium 140.115	59 <b>Pr</b> Praseodymium 140.908	60 <b>Nd</b> Neodymium 144.24	61 <b>Pm</b> Promethium 144.913	62 <b>Sm</b> Samarium 150.36	63 <b>Eu</b> Europium 151.966	64 <b>Gd</b> Gadolinium 157.25	65 <b>Tb</b> Terbium 158.925	66 <b>Dy</b> Dysprosium 162.50	67 <b>Ho</b> Holmium 164.930	68 <b>Er</b> Erbium 167.26	69 <b>Tm</b> Thulium 168.934	70 <b>Yb</b> Ytterbium 173.04	71 <b>Lu</b> Lutetium 174.967			
			89 <b>Ac</b> Actinium 227.028	90 <b>Th</b> Thorium 232.038	91 <b>Pa</b> Protactinium 231.036	92 <b>U</b> Uranium 238.029	93 <b>Np</b> Neptunium 237.048	94 <b>Pu</b> Plutonium 244.064	95 <b>Am</b> Americium 243.061	96 <b>Cm</b> Curium 247.070	97 <b>Bk</b> Berkelium 247.070	98 <b>Cf</b> Californium 251.080	99 <b>Es</b> Einsteinium [254]	100 <b>Fm</b> Fermium 257.095	101 <b>Md</b> Mendelevium 258.1	102 <b>No</b> Nobelium 259.101	103 <b>Lr</b> Lawrencium [262]			

<div style="background-color: #6a3d9a; color: white; padding: 5px; text-align: center;"> <b>21</b>  <span style="font-size: 2em; font-weight: bold;">Sc</span>                  Scandium                  44.95591             </div>	K 1s	4492 eV
	L <sub>1</sub> 2s	498 eV
	L <sub>2</sub> 2p <sub>1/2</sub>	403 eV
	L <sub>3</sub> 2p <sub>3/2</sub>	398 eV
<div style="background-color: #6a3d9a; color: white; padding: 5px; text-align: center;"> <b>92</b>  <span style="font-size: 2em; font-weight: bold;">U</span>                  Uranium                  238.03             </div>	K 1s	115606 eV
	L <sub>1</sub> 2s	21757 eV
	L <sub>2</sub> 2p <sub>1/2</sub>	20948 eV
	L <sub>3</sub> 2p <sub>3/2</sub>	17166 eV

- XAS is element specific: each element has characteristic absorption edge energies.
- With hard X-rays, we can probe many K and L edges across the periodic table.

# Typical layout of an XAS beamline

detectors

slits

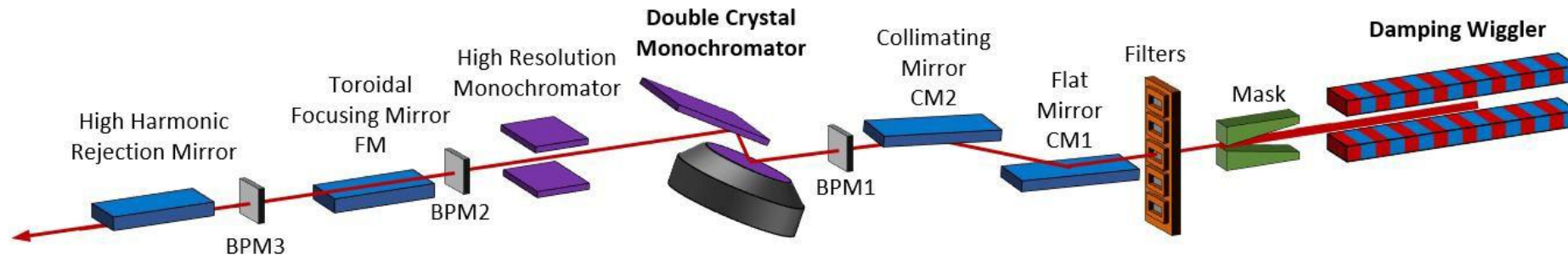
(mirror)

monochromator

(mirror)

source

← X-ray direction

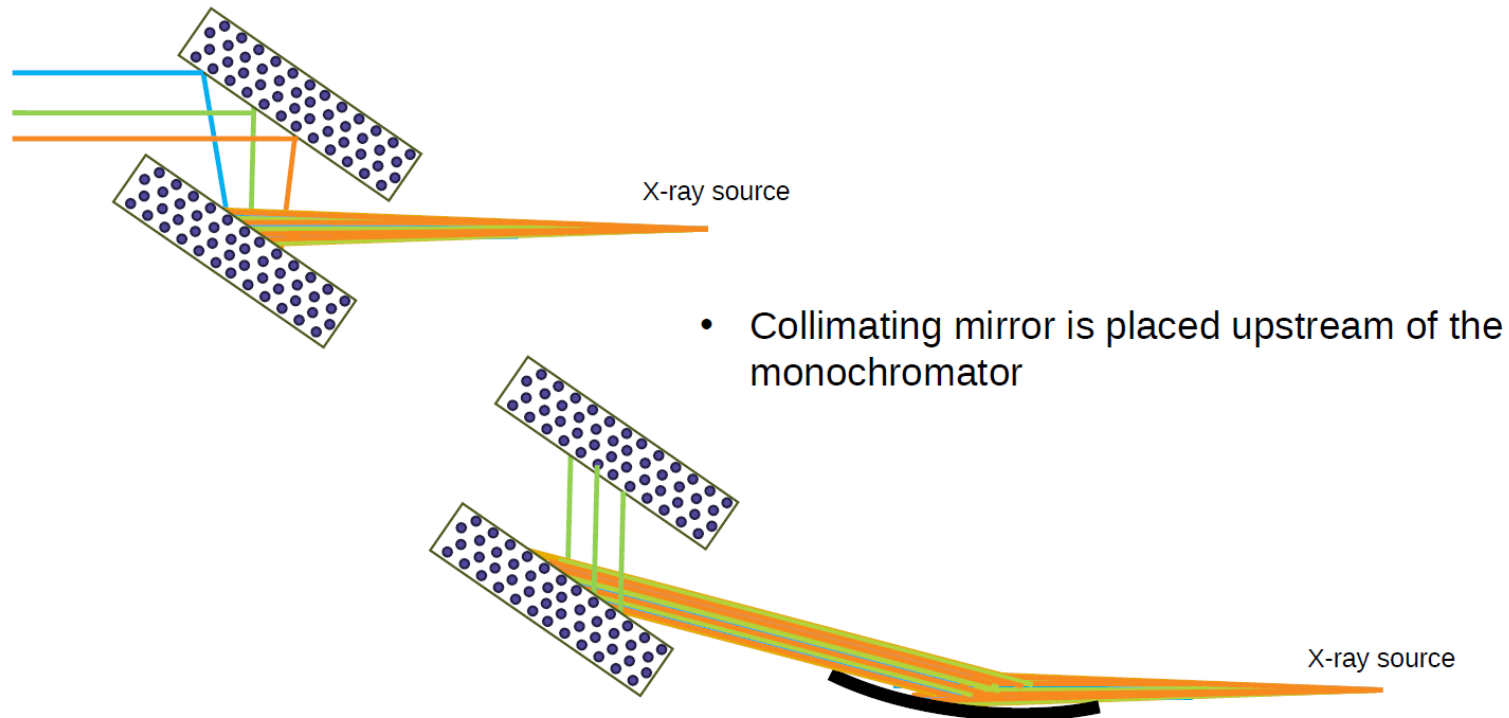


Simplest XAS beamline:  
• monochromator

Modern XAS beamline:  
• collimating mirror  
• monochromator  
• focusing mirror  
• harmonic rejection mirror

# Collimating mirror

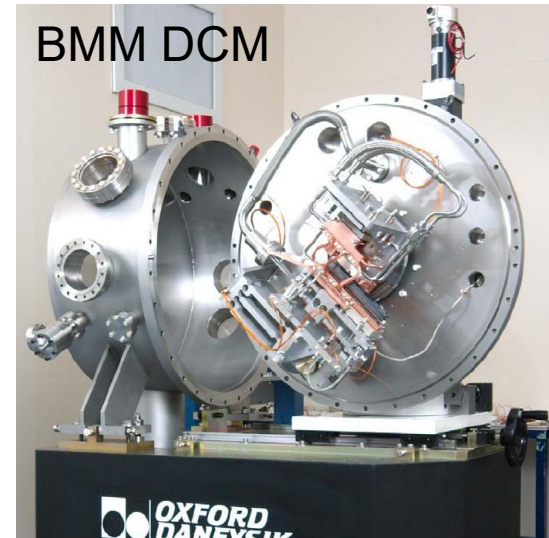
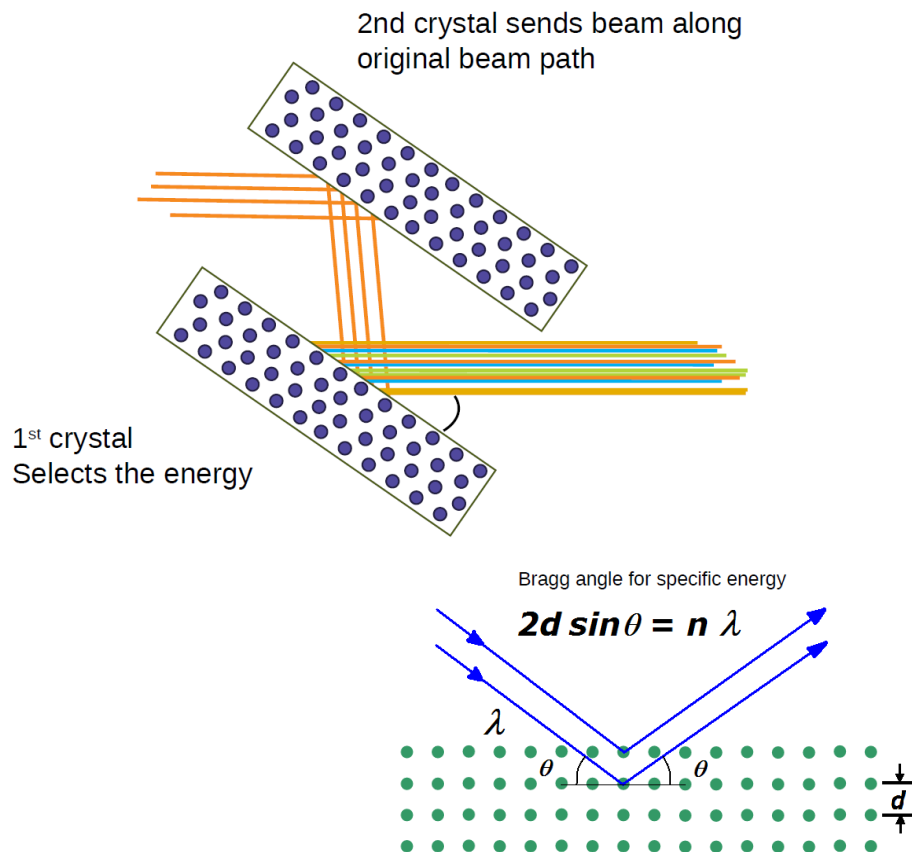
- Makes the divergent rays from the source more parallel.
- This helps ensure that X-rays hit the monochromator crystal at nearly the same angle, improving energy resolution.



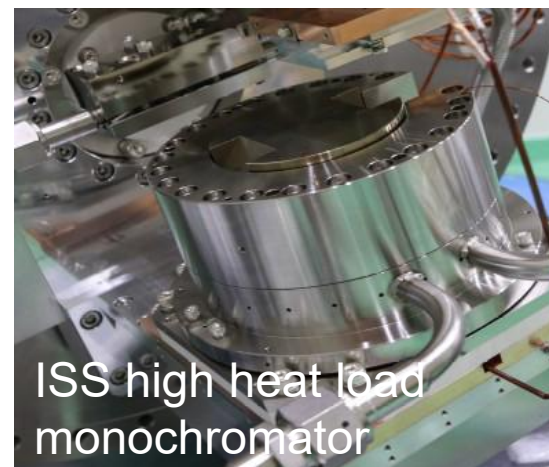
QAS collimating mirror

# Monochromator

- Uses Bragg diffraction from Si crystals to select a narrow energy bandwidth from the incoming beam.
- The X-ray energy is tuned by rotating the crystal angle.



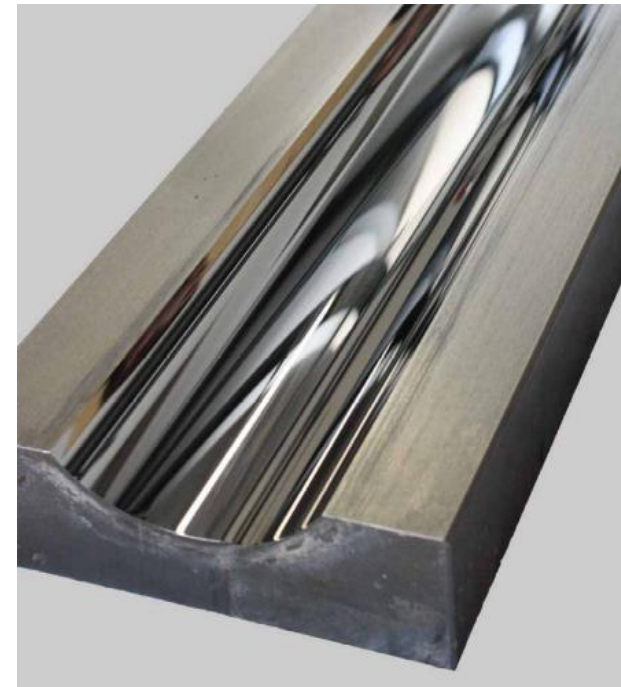
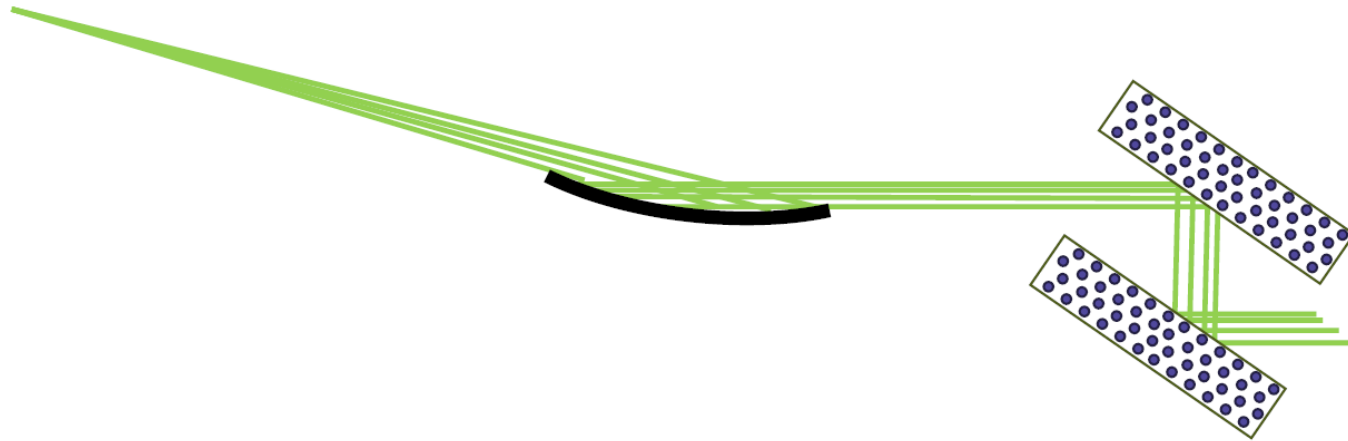
QAS Channel-Cut monochromator



# Focusing mirror

- Collimation mainly occurs vertically, while the beam continues to expand horizontally.
- The focusing mirror controls the beam size at the sample, from micrometers to millimeters depending on the application.

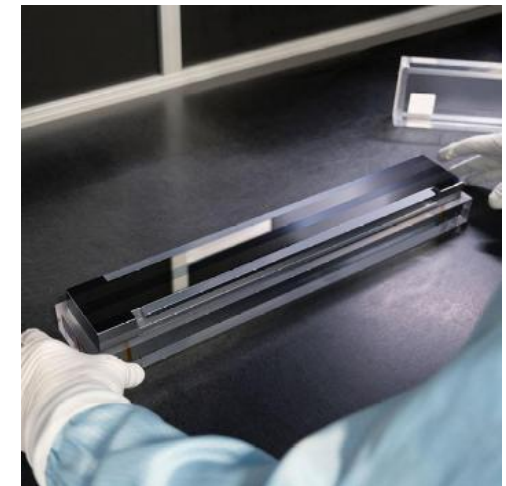
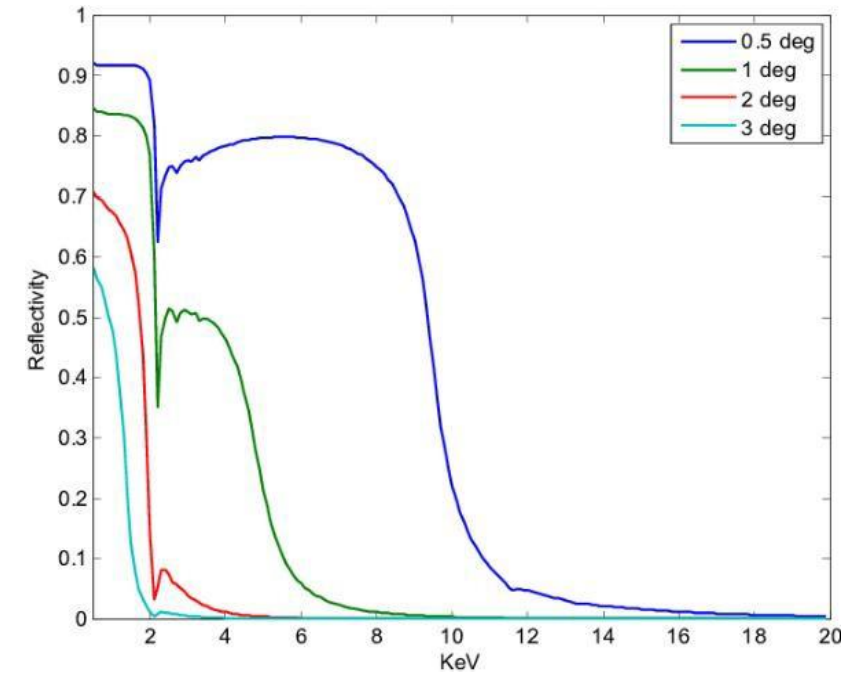
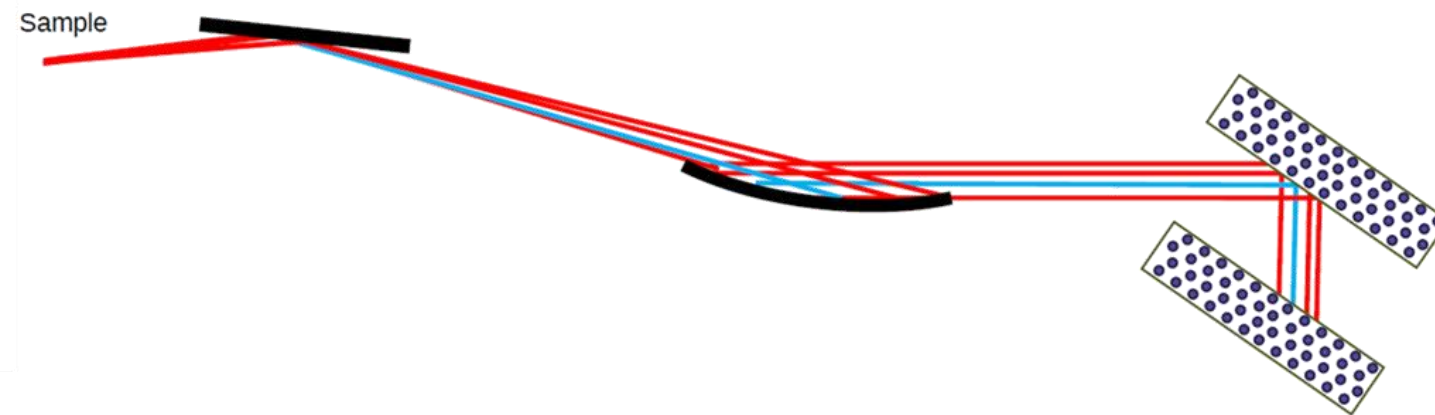
Sample



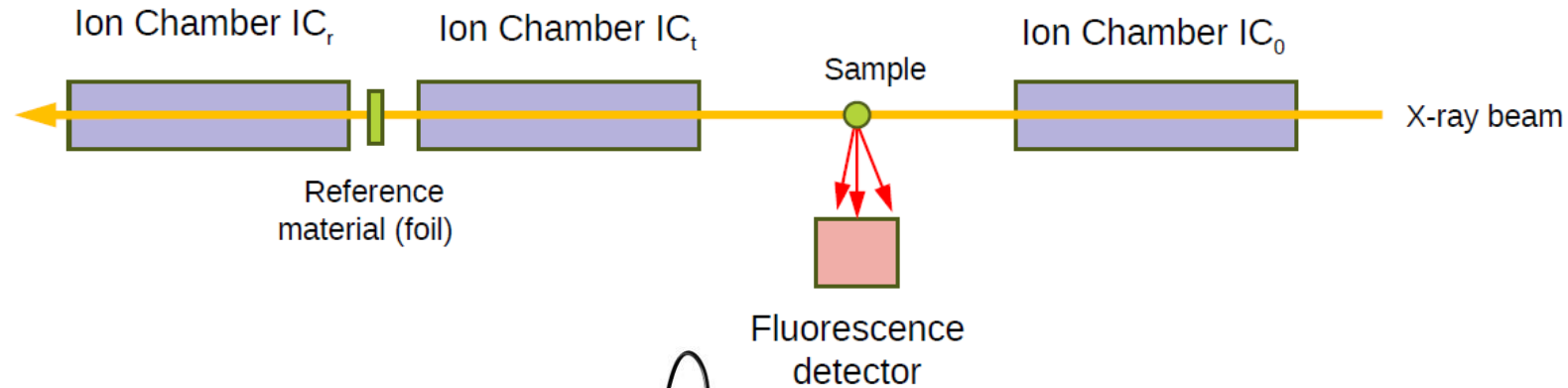
[https://crystal-scientific.com/mirror\\_cylindrical.html](https://crystal-scientific.com/mirror_cylindrical.html)

# Harmonic Rejection mirror

- Higher-order harmonics from the monochromator can contaminate the beam.
- A harmonic rejection mirror suppresses these unwanted higher-energy photons.



# Detectors for XAS



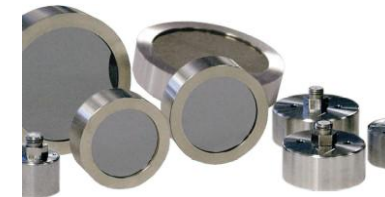
Detection mode:

- Transmission mode: ion chambers



<https://xds-oxford.com/products/>

- Fluorescence mode: Integrating or energy-discriminating fluorescence detectors

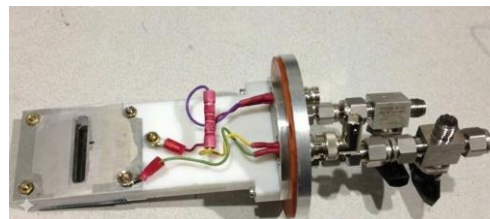


<https://www.mirion.com/products/technologies/>



<https://www.hitachi-hightech.com/us/en/products>

- Electron yield



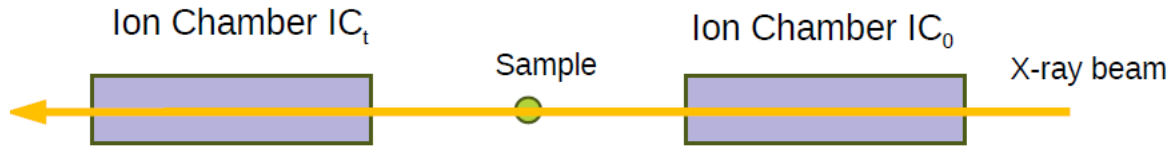
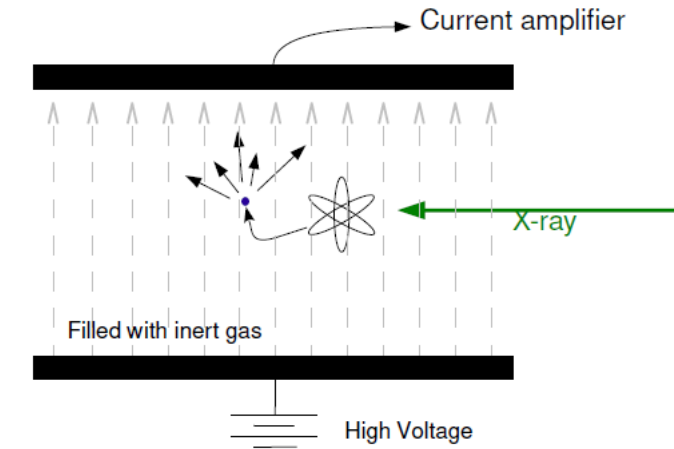
# Choosing the detection mode

- Transmission: concentrated, not too thick, edge step  $>0.1$ ,  $<2.0$
- Fluorescence: dilute samples:  $< 0.1$  absorption length edge step
- Electron yield: concentrated, thick or surface-sensitive measurements.

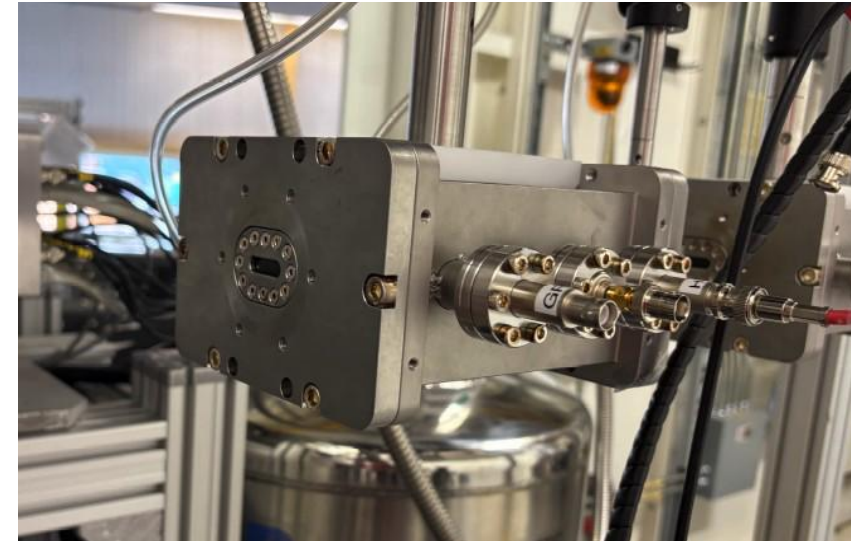
**Good sample preparation is essential for success!**

# Detectors - ion chambers

- Gas-filled detectors operated under high voltage (typically 500–2000 V).
- X-ray photons ionize the gas, generating a measurable current proportional to photon flux.



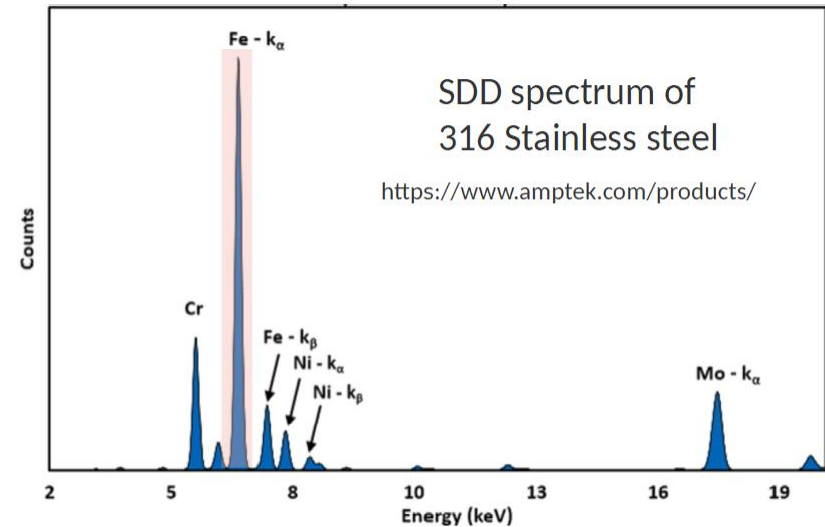
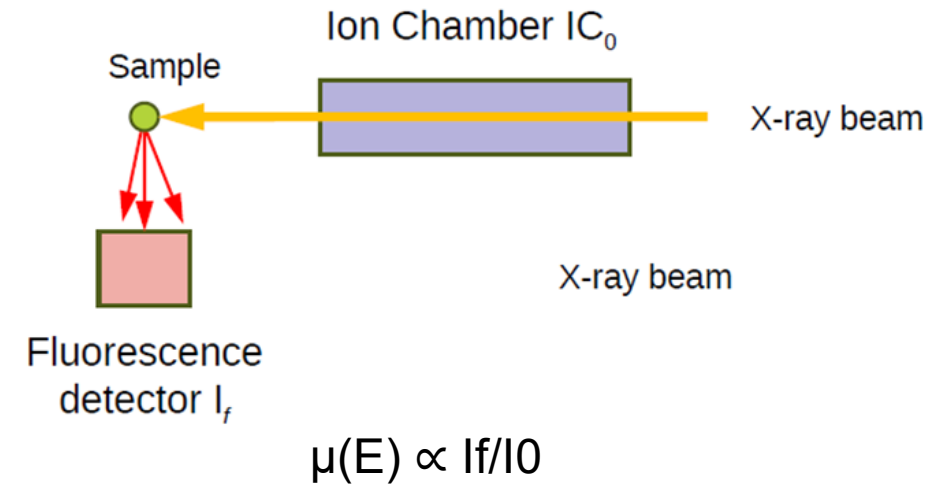
In transmission XAS:  $\mu(E) = \ln(I_0/I_t)$



# Detectors – Solid State detectors

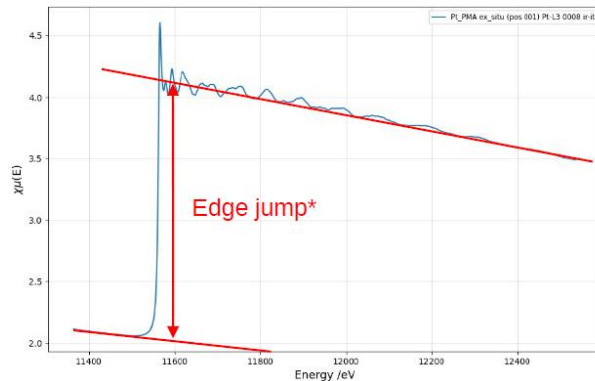
- They are commonly used for dilute samples and fluorescence-mode XAS.
- Typical materials:
  - Si
  - Ge
  - CdTe
- X-ray photons generate electron–hole pairs, and the collected charge is proportional to the detected fluorescence signal.

Maximum count rates of several hundred KHz total (signal+background)/channel

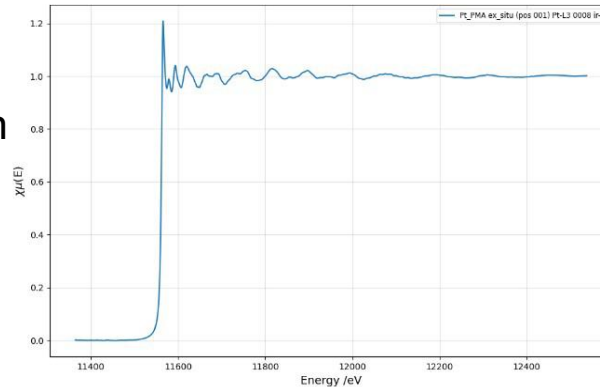


# Data Reduction

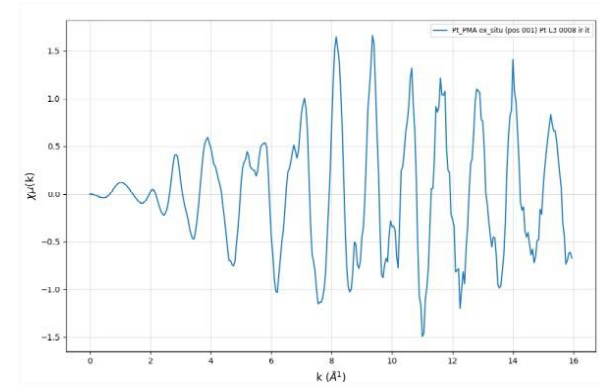
- Apply instrumental Corrections (e.g. detector dead-time)
- Normalize data to unit edge step
- Convert from  $E \rightarrow k$
- Subtract the background
- Apply k-weighting
- Fourier transform to R space
- Fourier Filter to isolate shells (optional)



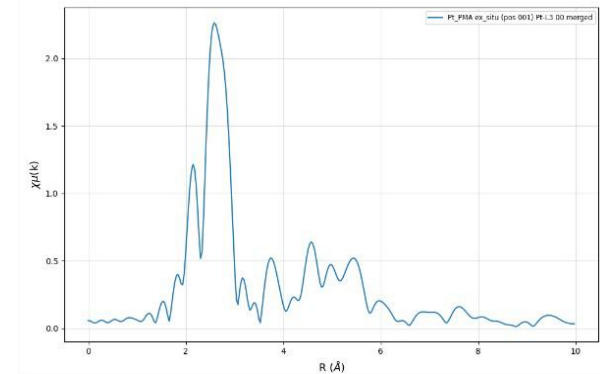
Normalization



$E \rightarrow k$



Fourier transform

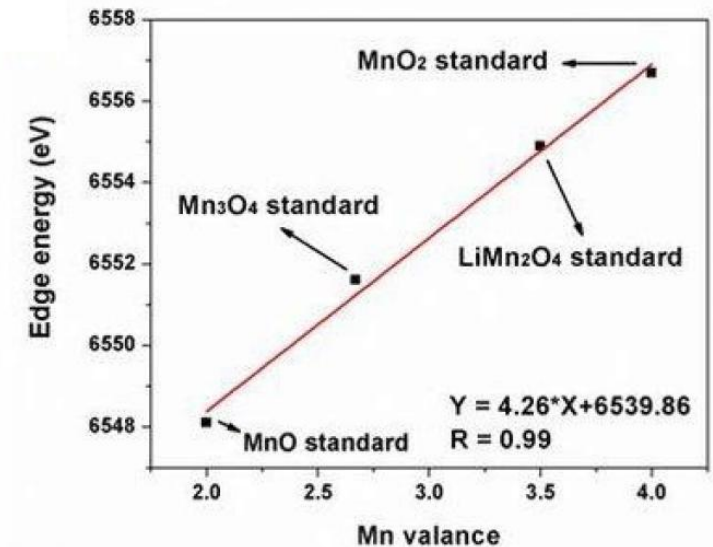
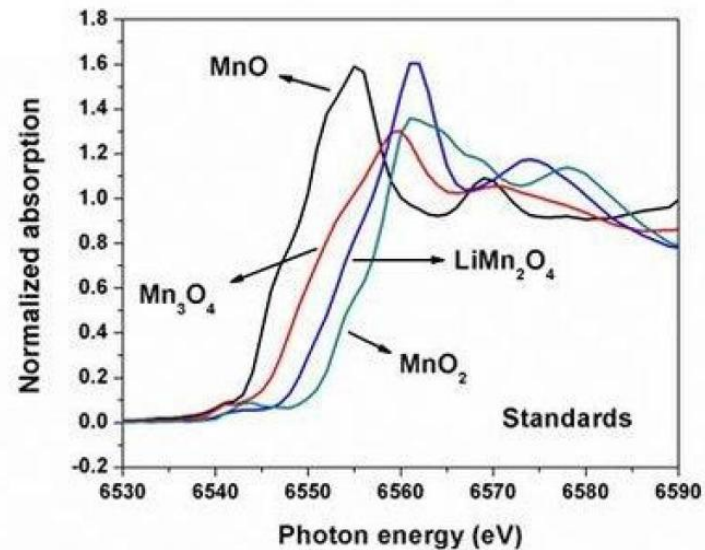


More in Lecture 2 at 11:10 am by Akhil Tayal:  
Data reduction and background removal

# What can we learn from XANES

## Oxidation and edge position

- The absorption edge shifts to higher energy as oxidation state increases due to weaker shielding effect of valence electrons and increase in Coulomb attraction.

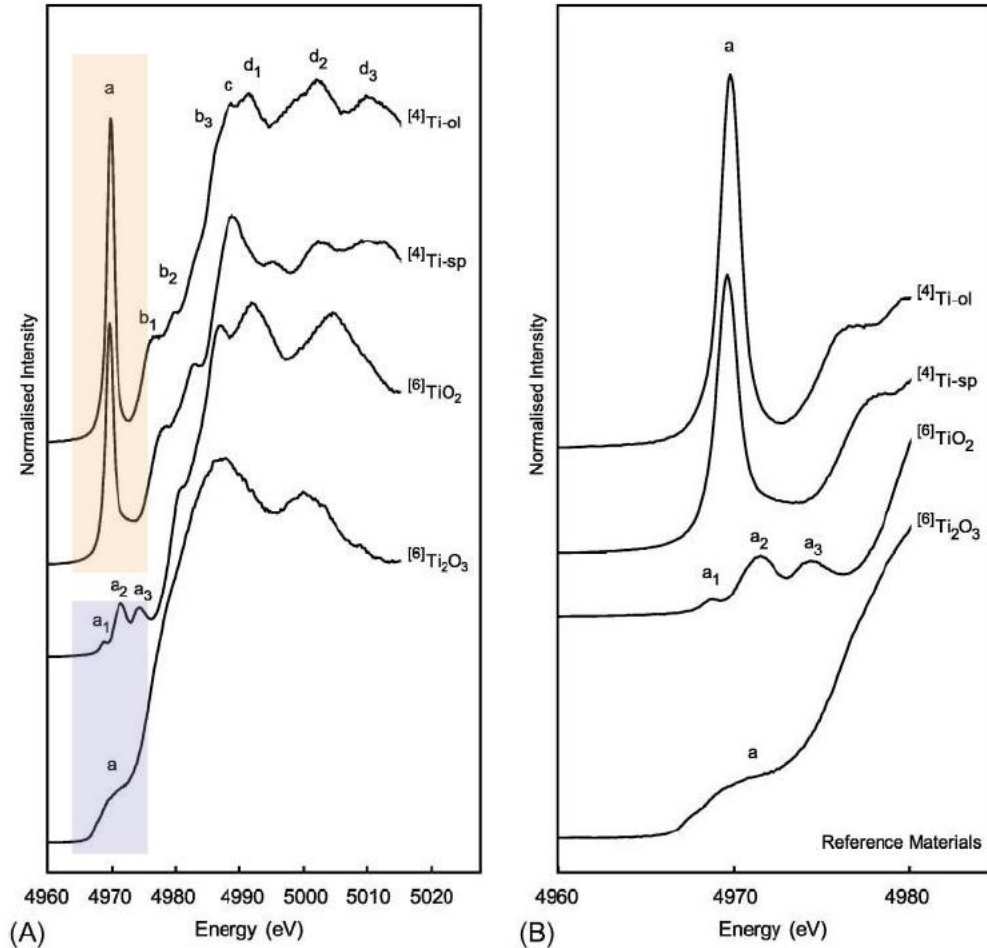


This example shows a linear relationship between edge position and effective oxidation state (it's not always linear).

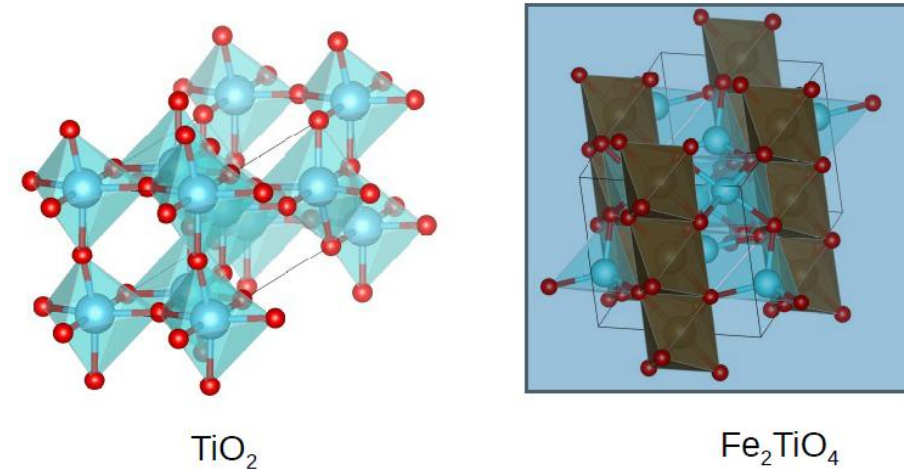
More in Lecture 3 by Jorge Moncada Vivas:  
XANES: fundamentals and analysis

# What can we learn from XANES

## Pre-edge – local symmetry



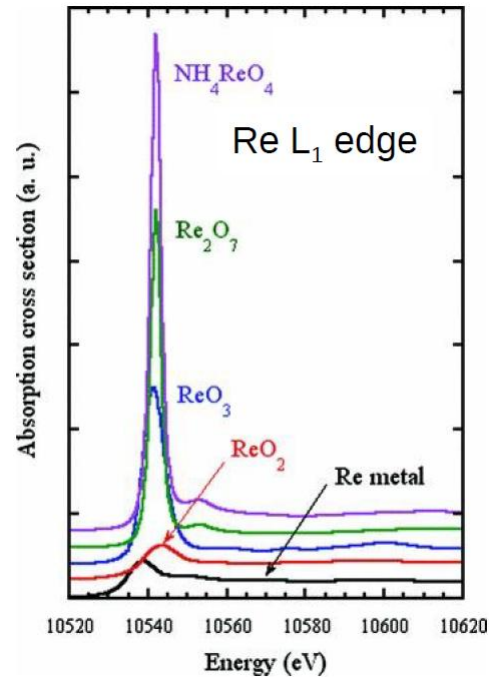
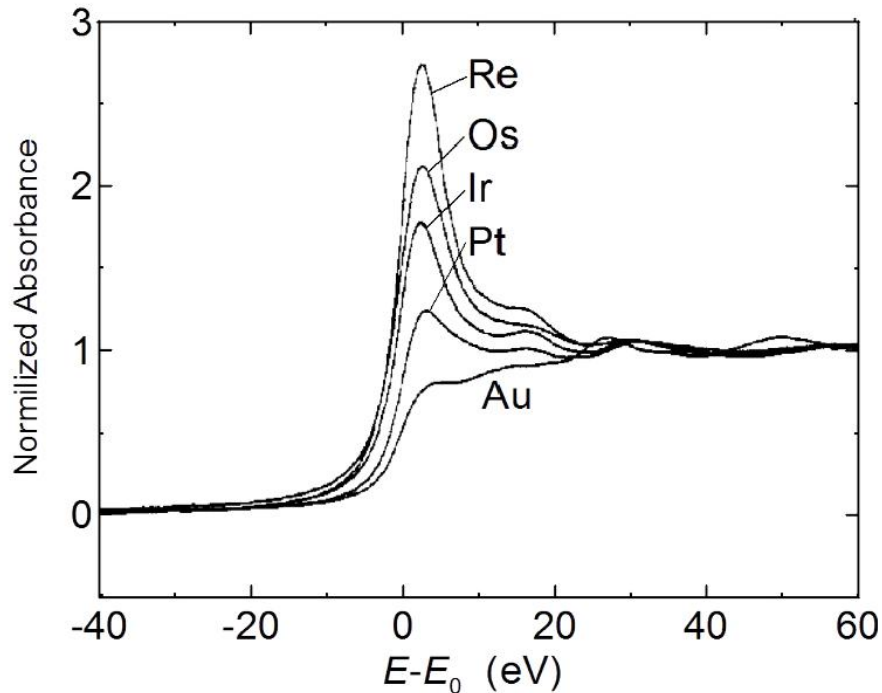
## Tetrahedral vs octahedral Ti



- In tetrahedral coordination, Ti lacks inversion symmetry, allowing 1s → 3d transitions to gain dipole intensity through p–d mixing.
- As a result, the pre-edge feature becomes stronger.

# What can we learn from XANES

White line intensity reflects the number of unoccupied d states.



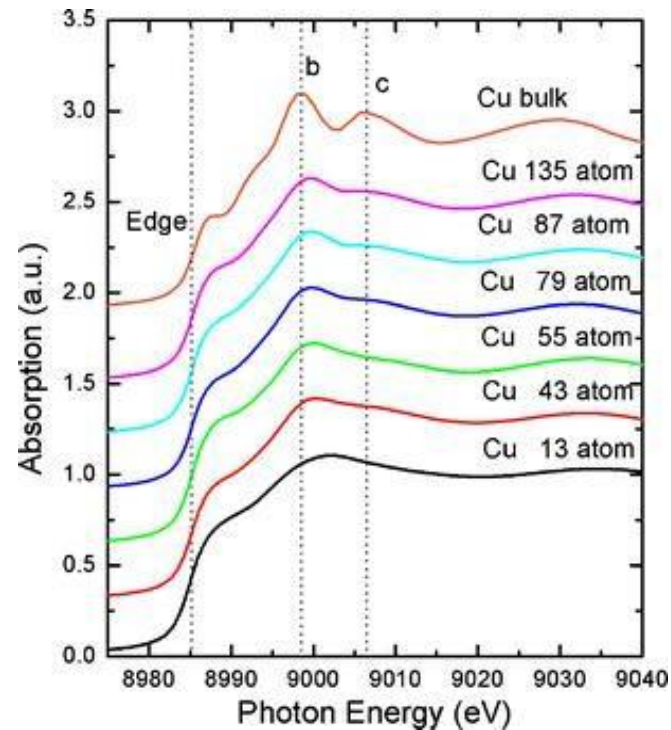
75 <b>Re</b> Rhenium 186.207	76 <b>Os</b> Osmium 190.23	77 <b>Ir</b> Iridium 192.22	78 <b>Pt</b> Platinum 195.08	79 <b>Au</b> Gold 196.967
$5d^5$	$5d^6$	$5d^7$	$5d^9$	$5d^{10}$

- For 5d elements, the white line arises from  $2p_{3/2} \rightarrow 5d$  transitions.
- White line intensity generally increases with the number of 5d holes.
- Gold shows little or no white line because excitation goes directly into the continuum.

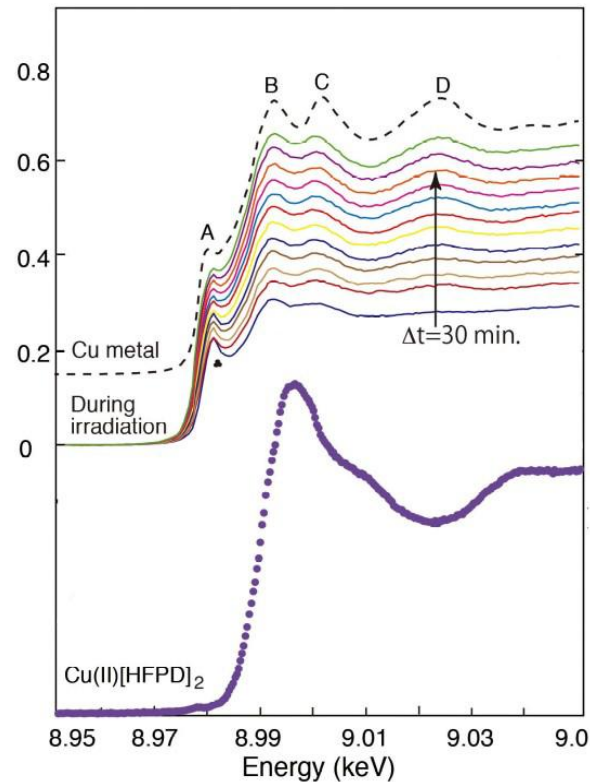
# What can we learn from XANES

## Nanoparticles vs bulk

Pronounced XANES features in bulk metals are often broadened or smeared in nanoparticles.



Analytical and Bioanalytical Chemistry  
399, 3033–3040 (2011)



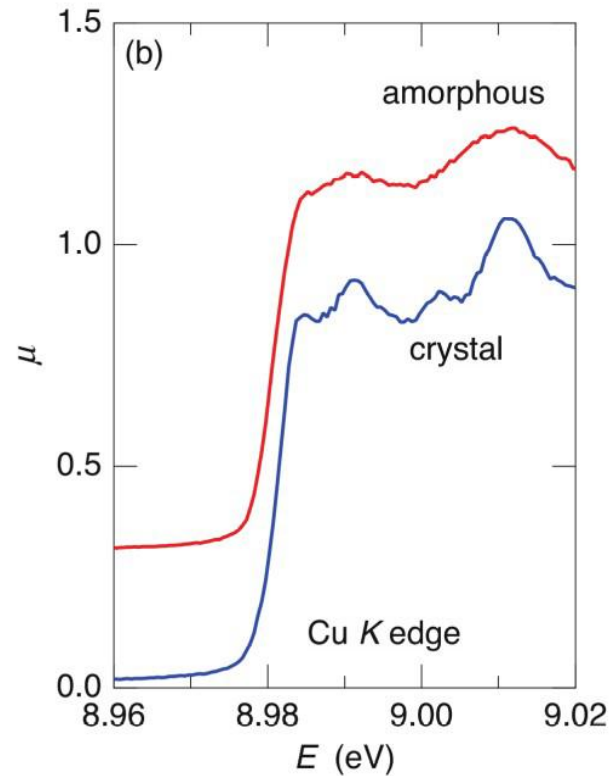
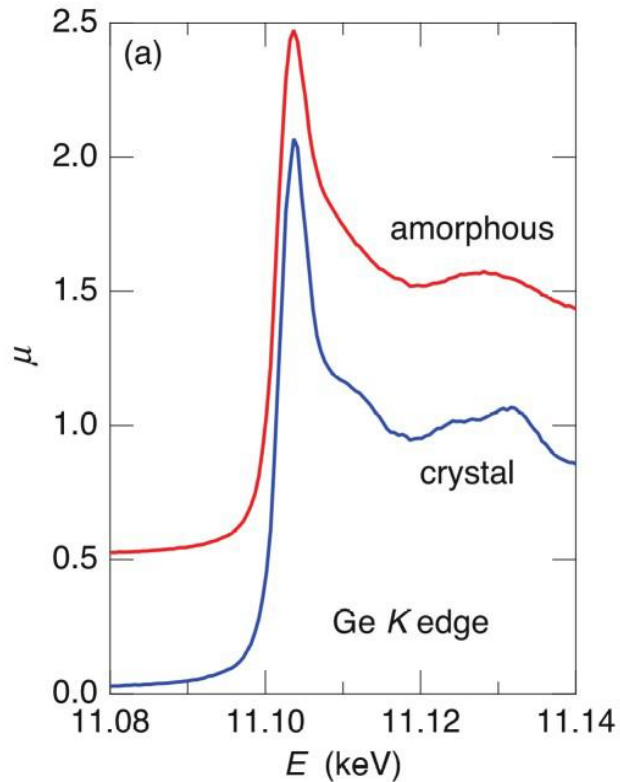
Scientific Reports 4, 7199 (2014)

Here the Cu particles grow from solution under radiolysis. As the size increases, the features in XANES sharpen

# What can we learn from XANES

## Amorphous vs. crystalline materials

Similar to the nanoparticle/bulk comparison, spectral features in crystalline materials are often sharper than in amorphous materials.

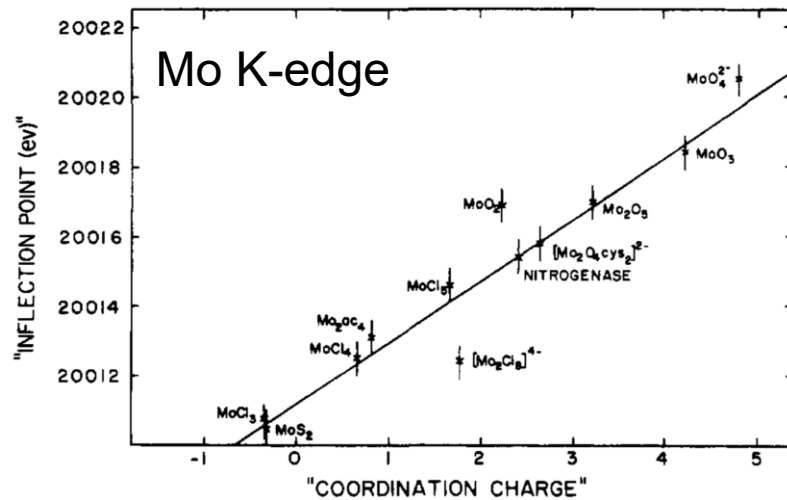


Phase-changing  
 $\text{GeCu}_2\text{Te}_3$  material  
before and after  
annealing

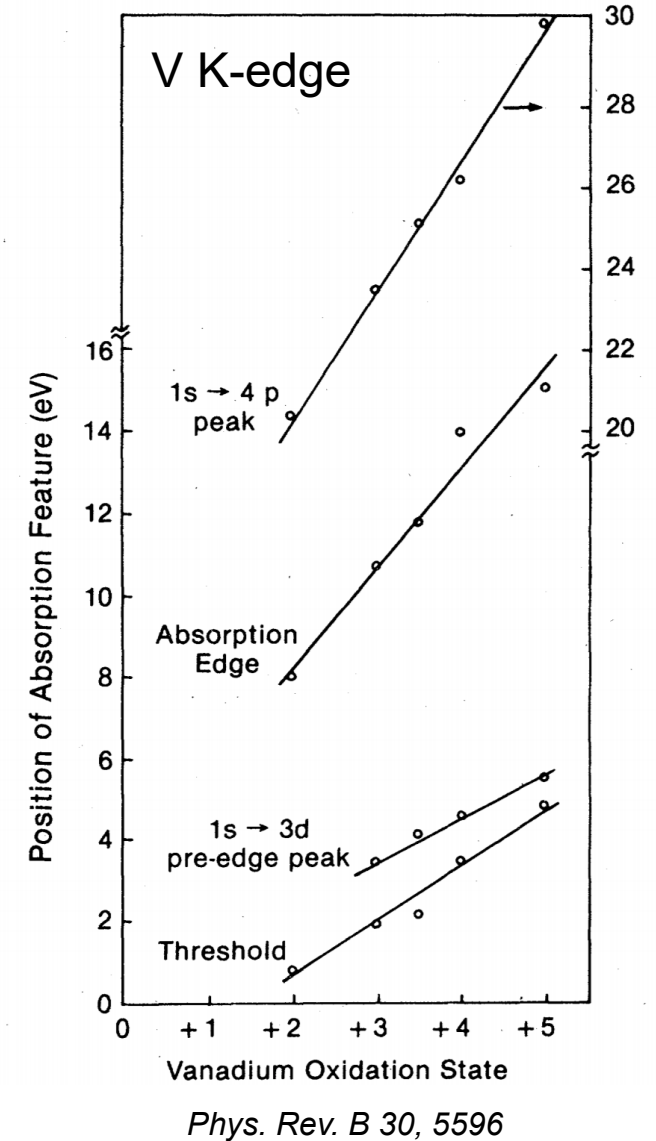
Journal of Optoelectronics and Advanced  
Materials 18, 248-253 (2016)

# Linear Combination Fitting

- Assumption: the XANES signature of a collection of atoms is the linear sum of the XANES from individual components
- Requirements:
  - appropriate reference standards
  - well-normalized spectra
  - consistent normalization



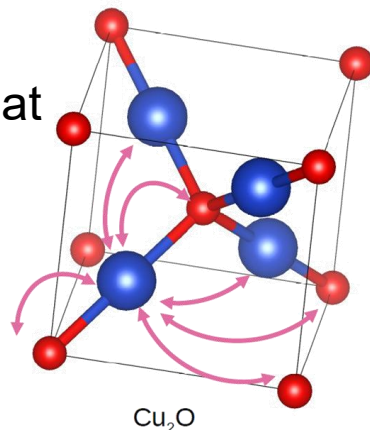
*J. Am. Chem. Soc.* 1976, 98, 5, 1287-1288



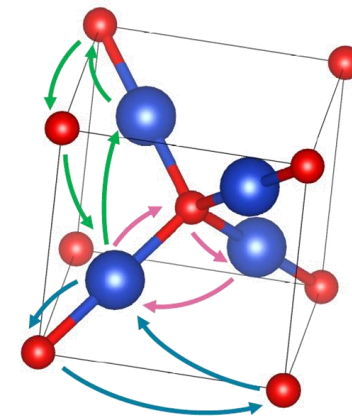
*Phys. Rev. B* 30, 5596

# How does a single path look like

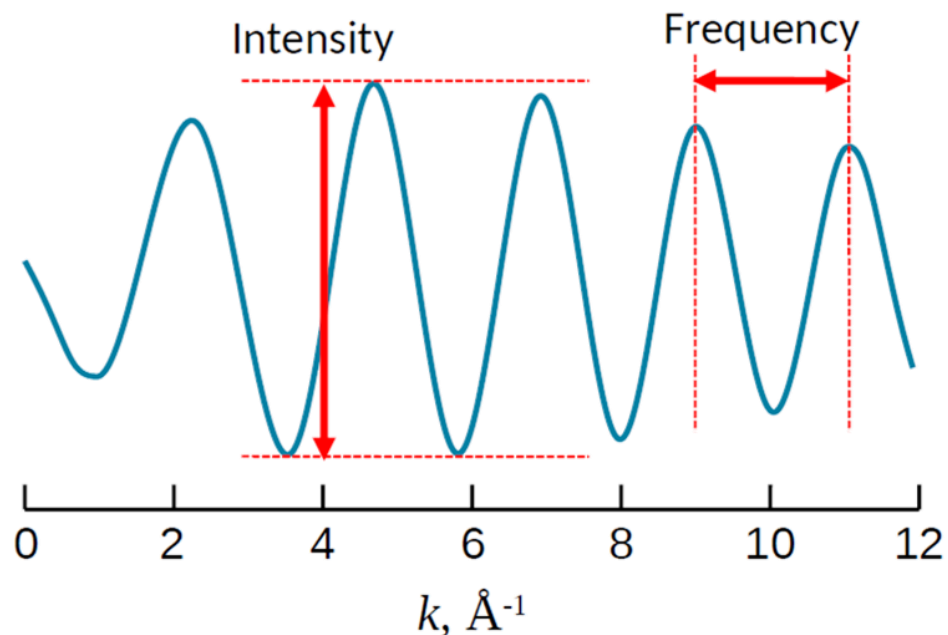
All "single" scattering paths are active at the same time



Scattering between more than two atoms - multiple scattering - plays a significant role



A single scattering path contributes a damped sinusoidal oscillation in k-space.



Intensity is defined by  
 $S_0^2, N_j, f_j(k), \sigma_j^2, R_j$

Frequency is defined by  
 $R_j, \delta_j(k)$

# EXAFS equation

$$\chi(k) = \sum_j \frac{N_j f_j(k) e^{-2k^2 \sigma_j^2}}{k R_j^2} \sin[2k R_j + \delta_j(k)] S_0^2$$

$N_j$  number of scattering atoms of type  $j$

$k$  photoelectron wavenumber,

$f_j(k)$  scattering amplitude

$R_j$  distance to scattering atoms

$\sigma_j^2$  Debye-Waller factor

$\delta_j(k)$  phase shift introduced by atom  $j$

$S_0^2$  amplitude reduction factor

$f_j(k)$  and  $\delta_j(k)$  are calculated (in this tutorial, using a code called FEFF\*)



\* To attempt an EXAFS fit you NEED to know something about the structure

We determine  $N$ ,  $R$  and  $\sigma_j^2$  from the fit\*\*



\*  $S_0^2$ ,  $N$  and  $\sigma_j^2$  are correlated

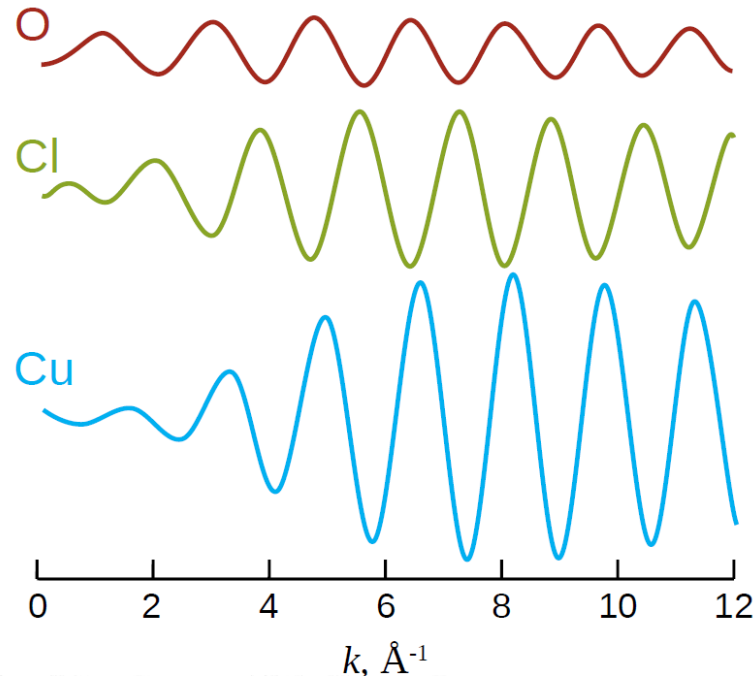
...and we sum all the scattering paths together!

More in Lecture 9 by Dali Yang:  
Understanding the EXAFS equation

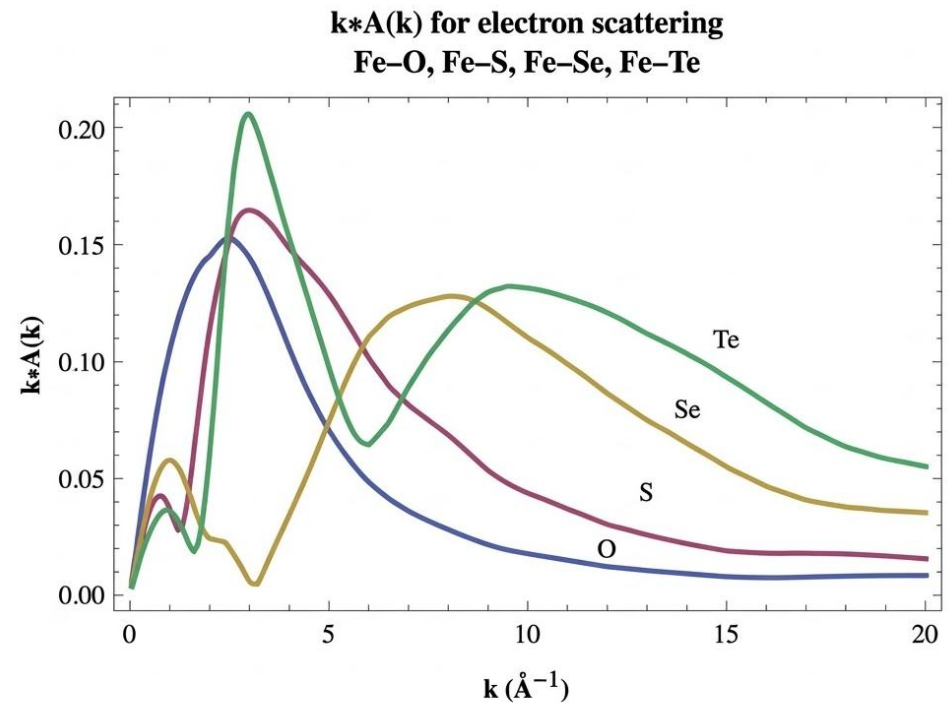
# Scattering Amplitudes

- The scattering amplitude  $f_j(k)$  depends on the identity of the neighboring atom. Heavier neighboring atoms generally scatter more strongly.

$$\chi(k) = \sum_j \frac{N_j f_j(k) e^{-2k^2 \sigma_j^2}}{k R_j^2} \sin[2k R_j + \delta_j(k)] S_0^2$$

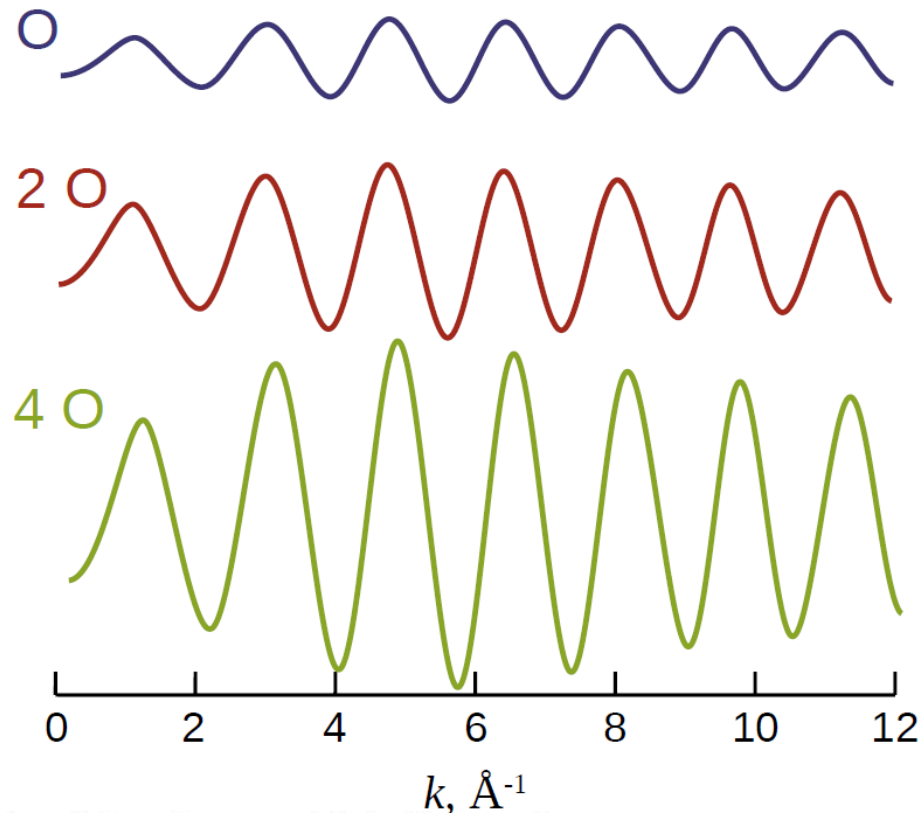


In this illustration, O, Cl, and Cu are shown at the same distance from the absorber for comparison purposes.



# Number of scatterers

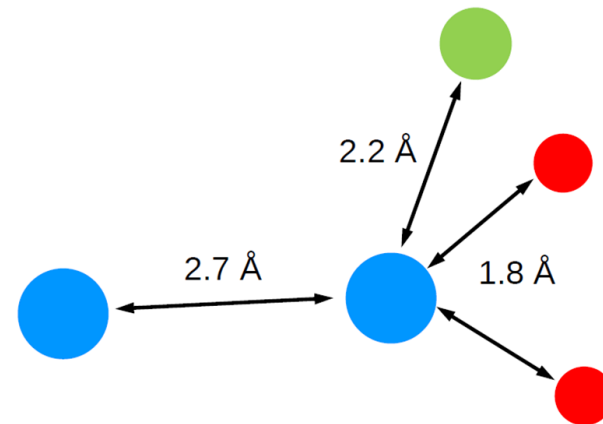
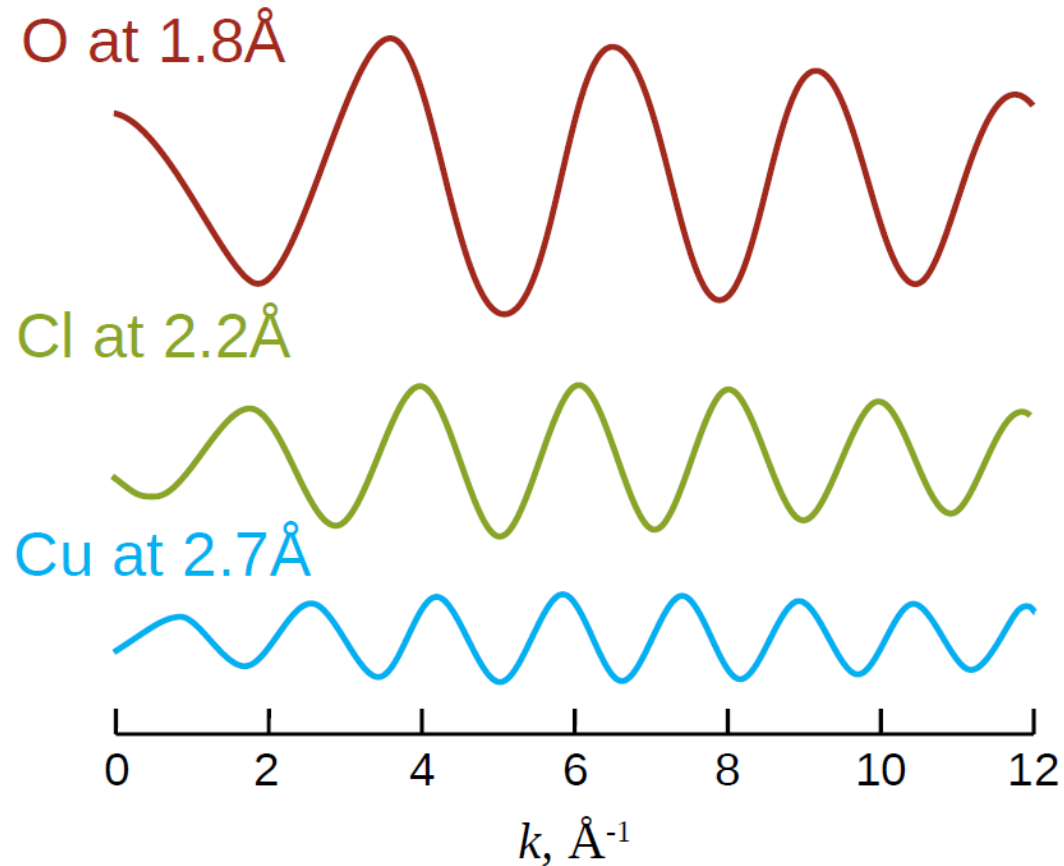
Larger coordination number increases EXAFS amplitude.



$$\chi(k) = \sum_j \frac{N_j f_j(k) e^{-2k^2 \sigma_j^2}}{k R_j^2} \sin[2k R_j + \delta_j(k)] S_0^2$$

# Distance to scatterers

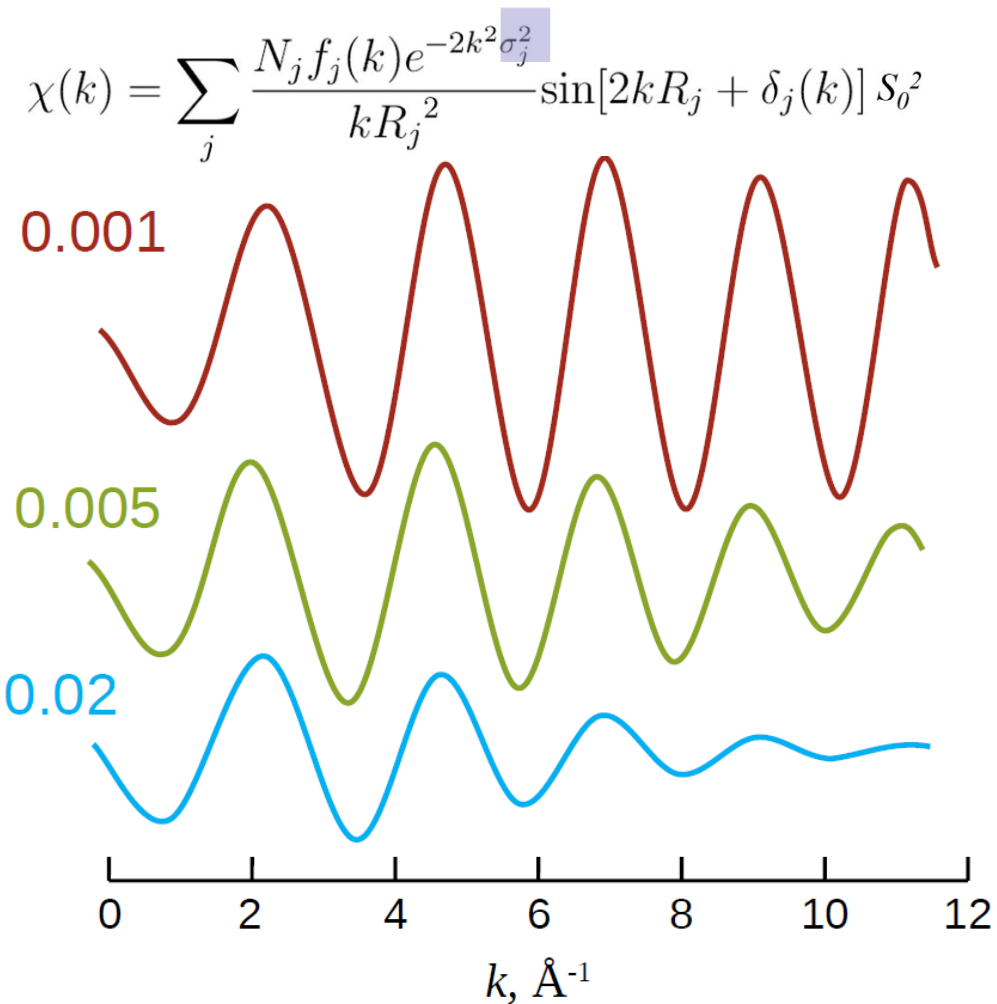
As the scatterer moves farther away, the oscillation frequency increases while the amplitude decreases.



$$\chi(k) = \sum_j \frac{N_j f_j(k) e^{-2k^2 \sigma_j^2}}{k R_j^2} \sin[2k R_j + \delta_j(k)] S_0^2$$

# Debye-Waller factor

The Debye–Waller factor reflects disorder, including thermal and structural disorder.



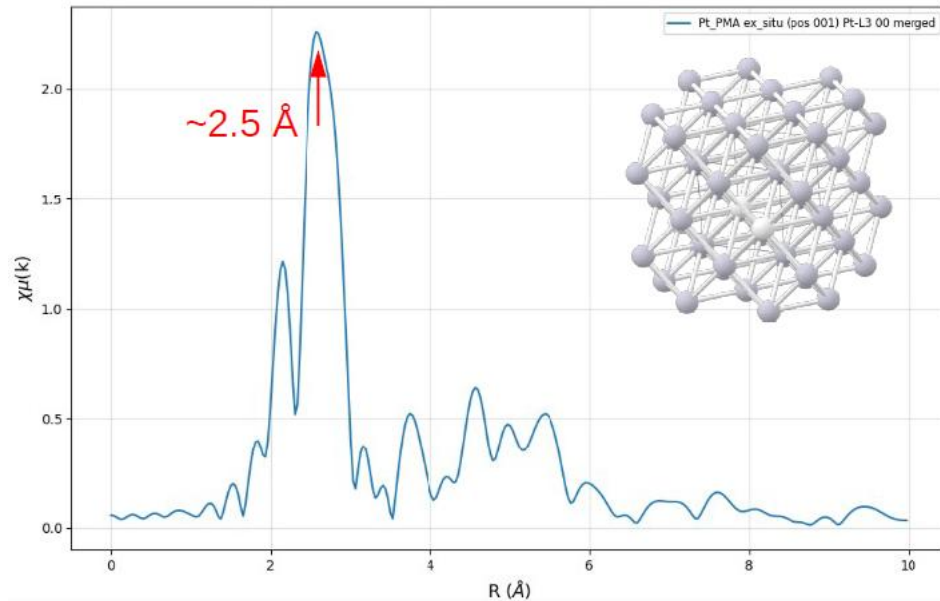
## Difference of Debye-Waller factor in XRD vs EXAFS

- In XRD: describes atomic displacement from lattice positions
- It reflects:
  - thermal vibration
  - positional disorder in the crystal lattice
- In EXAFS: describes fluctuations in the distance between the absorber and neighboring atoms
- it includes:
  - thermal vibration
  - structural disorder
  - distribution of bond lengths

EXAFS Debye–Waller factor is usually larger because it includes relative motion of two atoms.

# Phase shift

Peaks in the Fourier transform do not appear exactly at the true interatomic distances.



The real Pt distance is 2.79 Å

$$\chi(k) = \sum_j \frac{N_j f_j(k) e^{-2k^2 \sigma_j^2}}{k R_j^2} \sin[2k R_j + \delta_j(k)] S_0^2$$

$\delta_j(k)$  is an element specific phase shift that causes the distance in FT-EXAFS to underestimate the real one by up to 0.5 Å.

# Amplitude reduction factor $S_0^2$

$S_0^2$  accounts for many-body amplitude reduction and is usually treated as an empirical fitting parameter.

$$\chi(k) = \sum_j \frac{N_j f_j(k) e^{-2k^2 \sigma_j^2}}{k R_j^2} \sin[2k R_j + \delta_j(k)] S_0^2$$

It is typically determined using a well-defined reference sample, such as a metal foil.

$$\chi(k) = \sum_j \frac{N_j f_j(k) e^{-2k^2 \sigma_j^2}}{k R_j^2} \sin[2k R_j + \delta_j(k)] S_0^2$$

$N_j$  and  $S_0^2$  are correlated, so an incorrect  $S_0^2$  leads to inaccurate coordination numbers.

More in Lecture 4 and 8 by Bruce Ravel:  
EXAFS analysis I & II

# What XAF spectroscopy can do — Key Takeaways

## 1. Element-specific probing

- Individual elements in complex or heterogeneous materials can be studied independently.
- All atoms of the selected element contribute — there are **no spectroscopically silent atoms**. However, very light atoms such as hydrogen scatter the photoelectron weakly and are often difficult to detect.

## 2. Electronic structure information: oxidation state and orbital occupancy can often be determined from **XANES**

## 3. Local structural information: precise information on **bond distances, coordination numbers, neighboring atom types** ( $Z$ can usually be determined to $\pm 5$ ), and **disorder**

## 4. Applicable to **crystalline and non-crystalline systems**

## 5. Works for **any phase of matter**: gas, liquid, or solid

## 6. **Works in realistic environments**: Measurements can be performed **in situ / operando**, under conditions close to the natural or working states.

# Reference

